

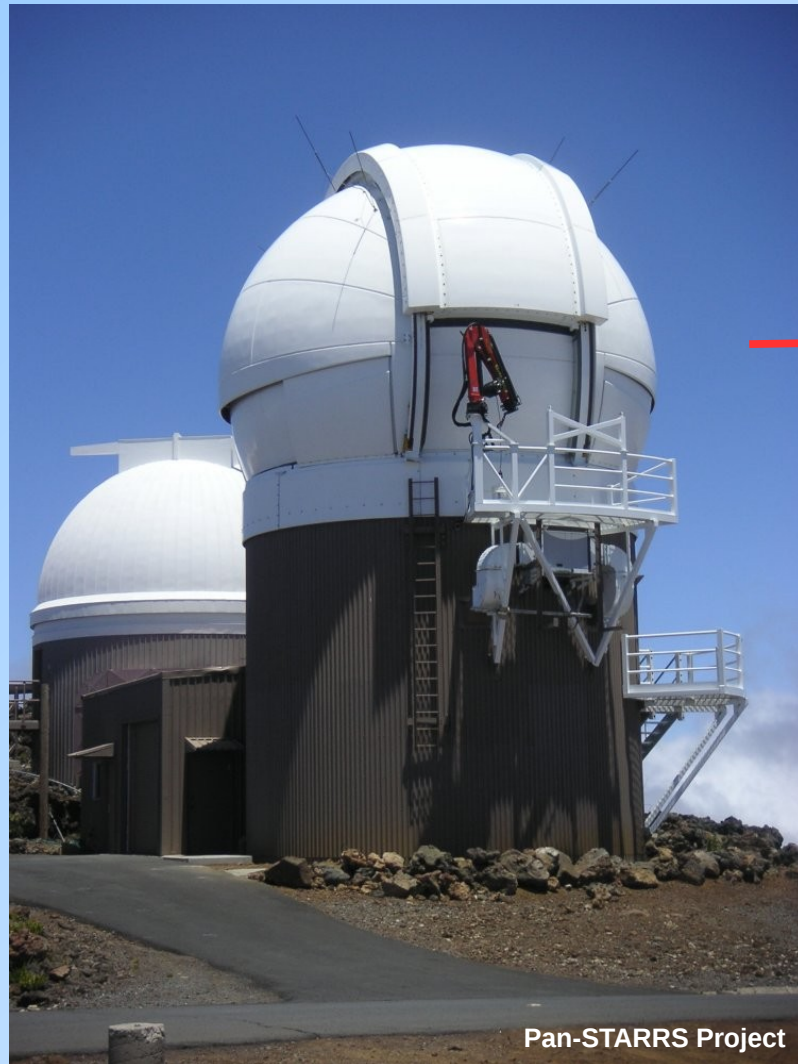
Precision Astrometry and Photometry from Pan-STARRS 1



Eugene Magnier

thanks to
John Tonry
Doug Finkbeiner

Pan-STARRS 1: a 1.8m survey telescope (1.4Gpix & 7deg² F.O.V.)



PS1 Survey Mission:
5/2010 – 3/2014

Google

Surveys:

3pi : ~12 / filter

MD : ~500 / filter

others : M31, Solar
System, STS



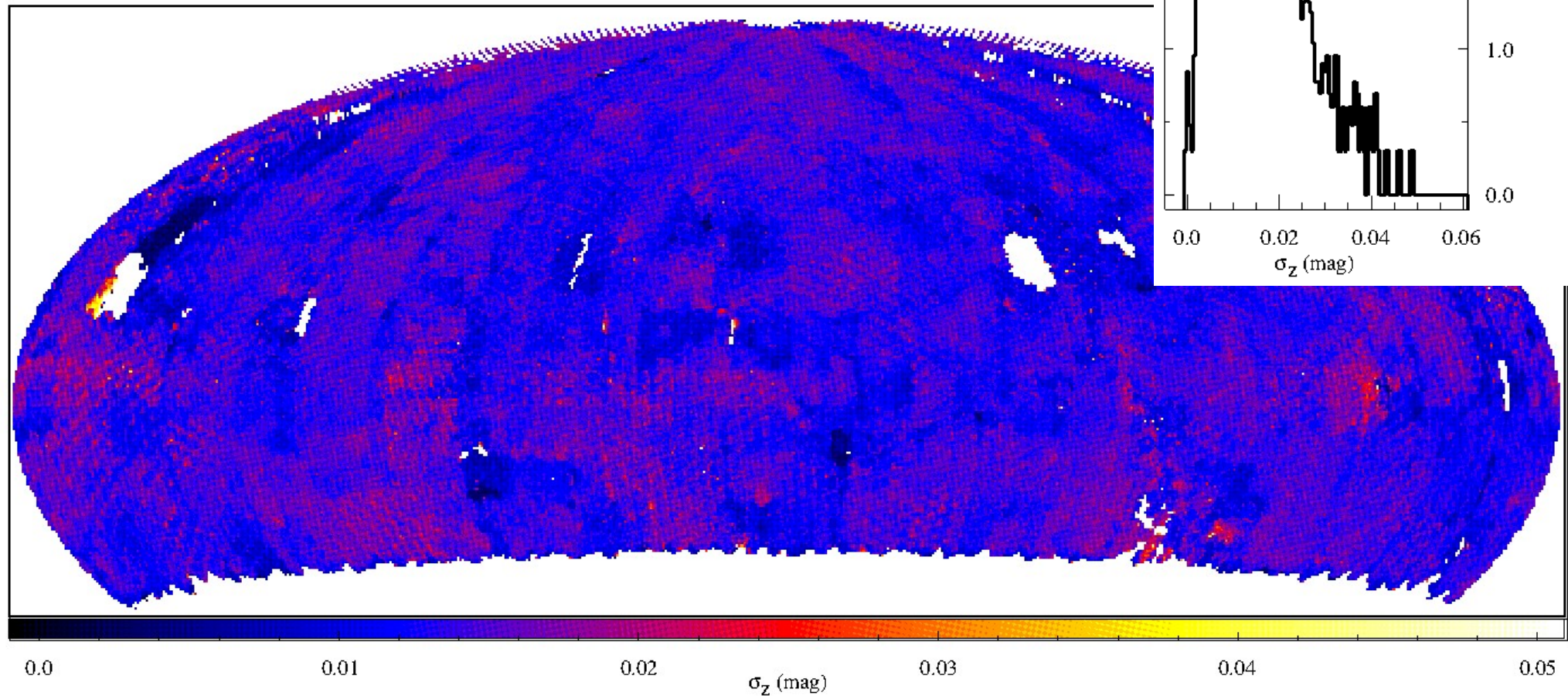
PS-1 on Haleakala

PS1 consortium members



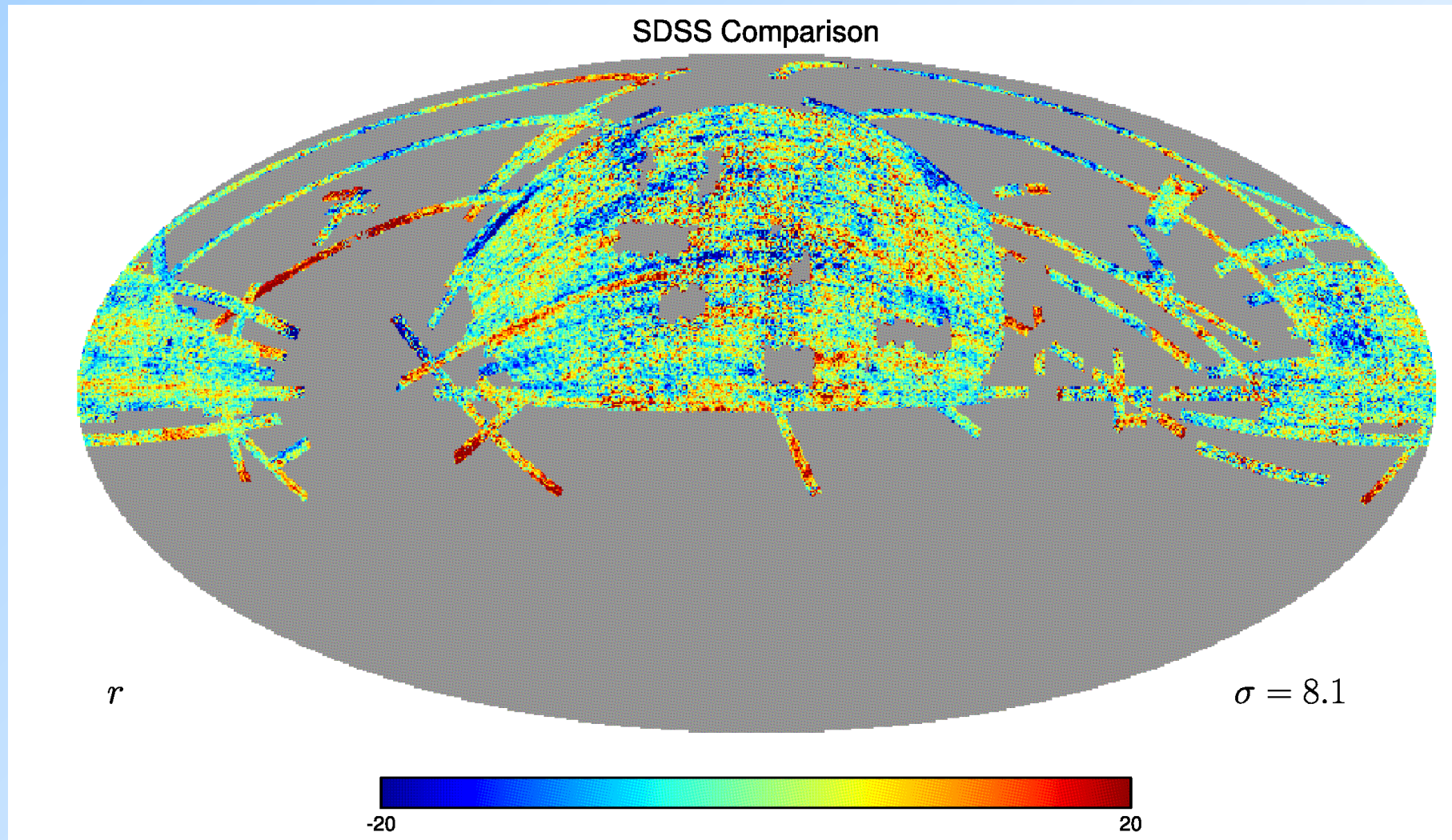
Photometry Performance

- 3pi reliability : (*grizy*) = (8, 7, 9, 11, 12) mmag
- MD reliability : all filters < 6 mmag
- per-exposure scatter ~ 10 – 15 millimags



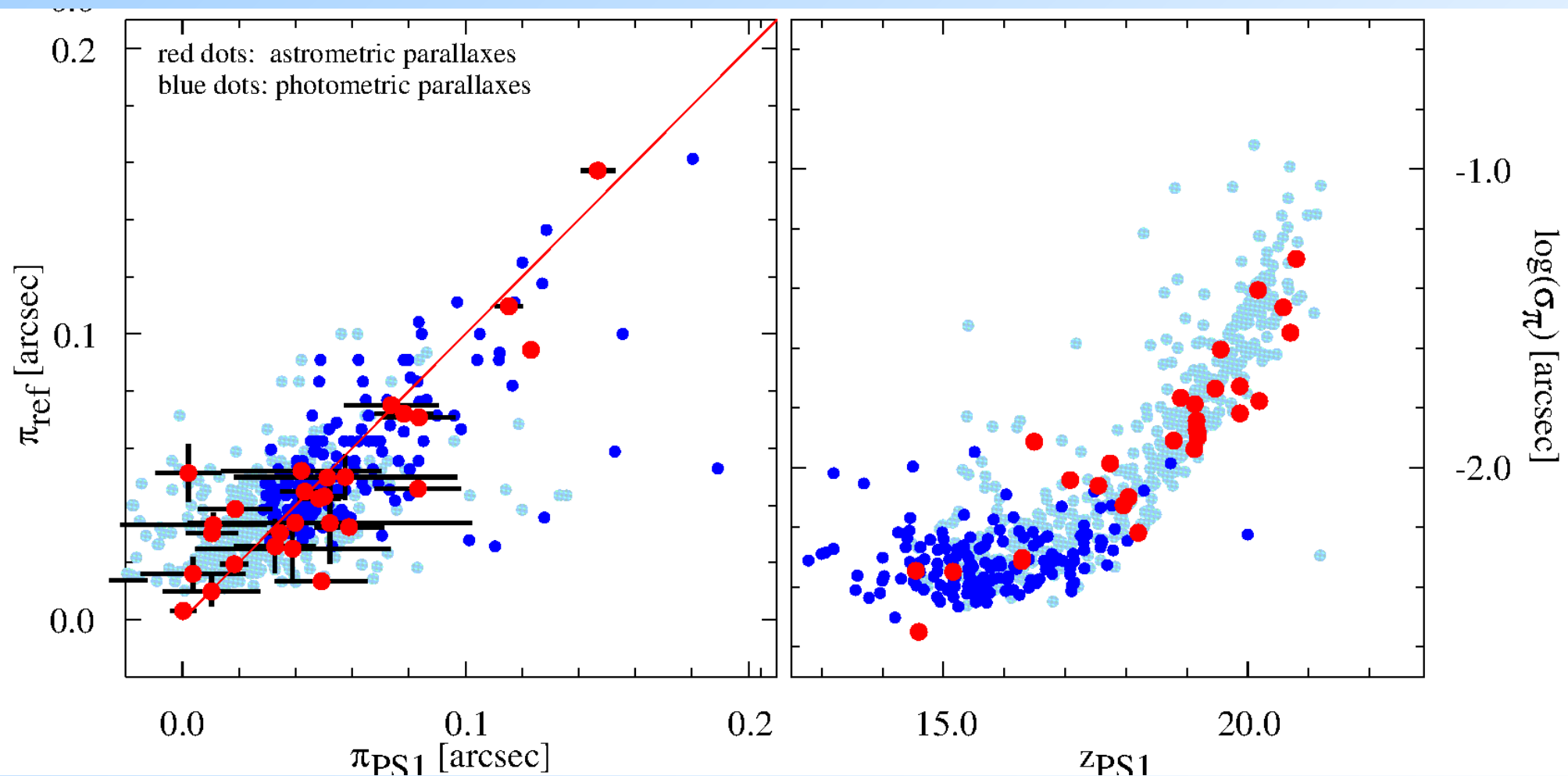
Photometry Performance

- PS1 vs SDSS shows systematic SDSS structures
 - Doug Finkbeiner is working on SDSS recalibration with PS1
 - eBoss will use PS1-based recalibration for target selections
 - In discussion for inclusion in future SDSS DR



Astrometry Performance

- Per-detection : 18 - 20 mas (1D, depends on chip)
- Parallax limit (1.5yr) : 3-4 mas
- Proper-motion limit (1.5yr) : ~5 mas/yr

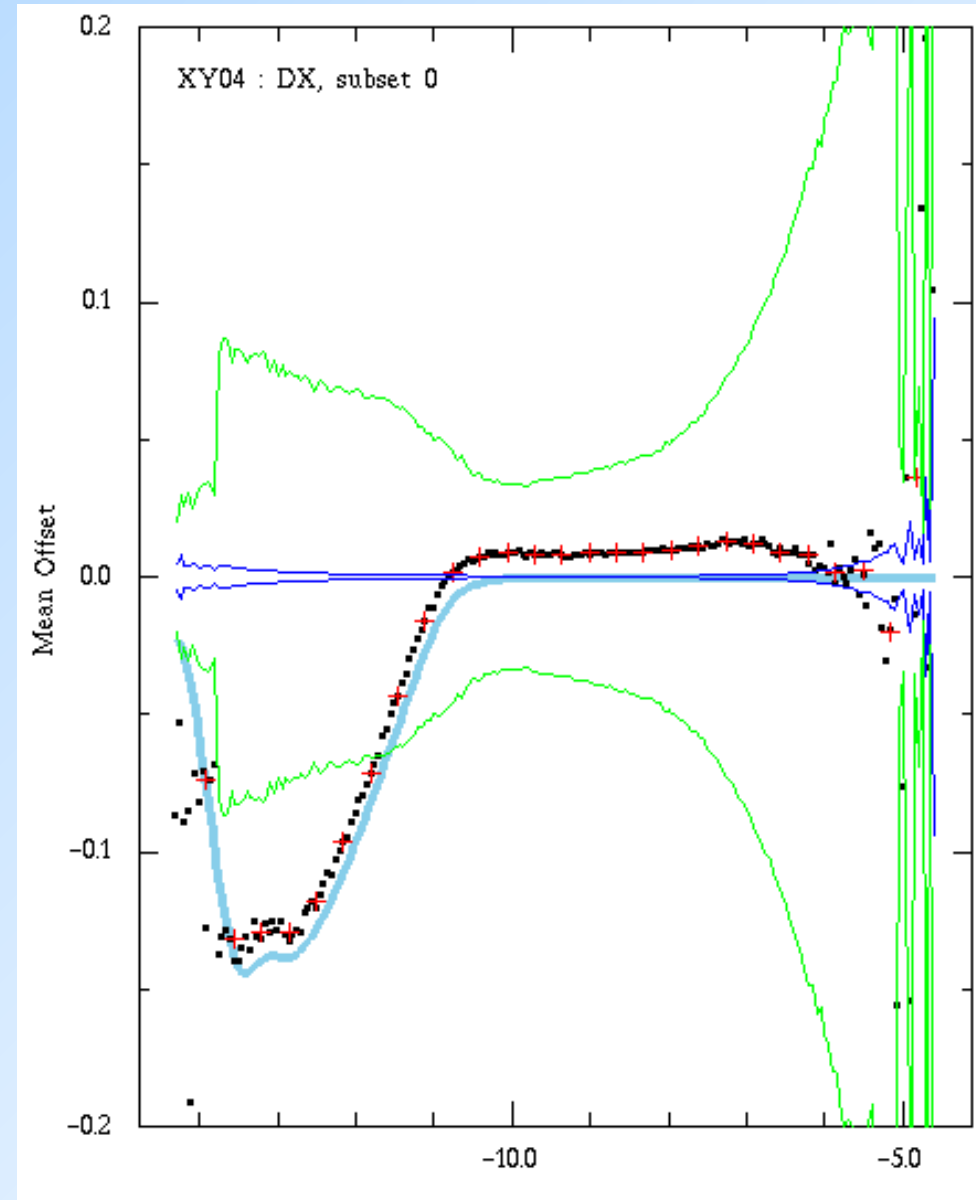


Instrumental Effects

- Bias, Dark, Flat
- Stellar photometric flat
- non-linear darks
- video-dependent dark structure
- persistence
- corner glows
- poor CTE regions
- non-linear response (per cell) at faint end
- non-linear response at bright end
- cross-talk (interchip and intrachip)
- row-by-row bias variations
- 2D PSF variations
- ghosts
- glints

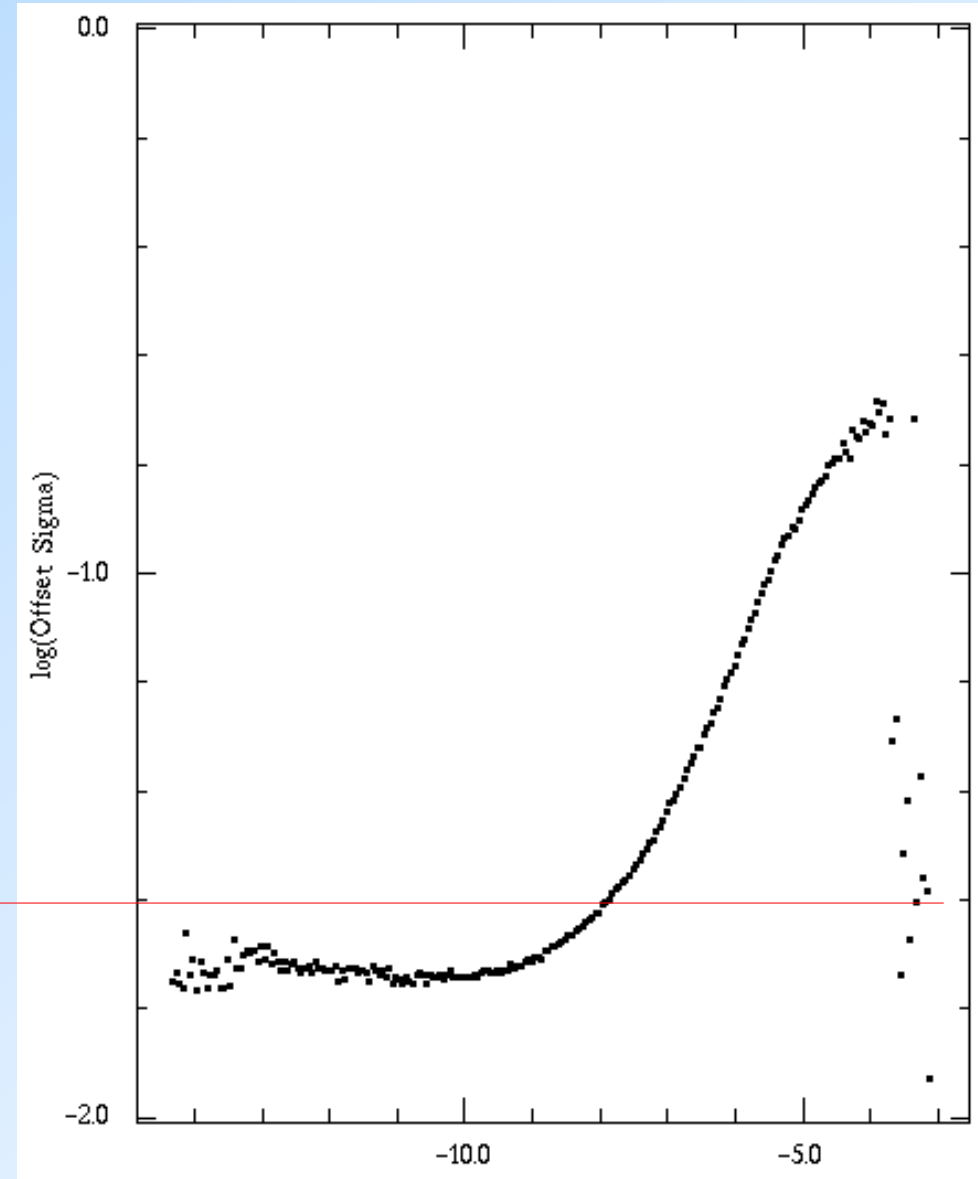
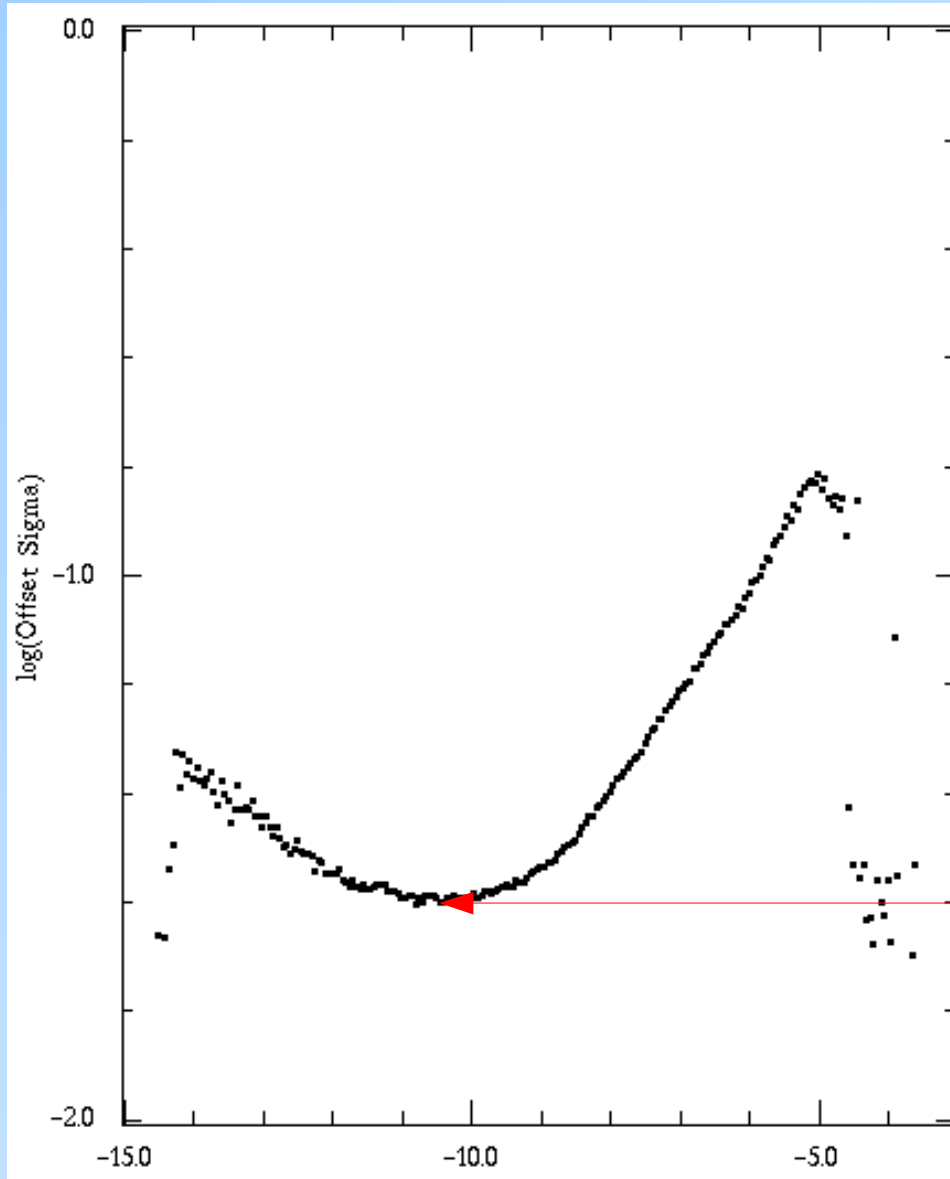
The Koppenhoefer Effect

- bright-end position bias as function of magnitude
- discovered by Johannes Koppenhoefer in STS data
- camera voltages adjusted May 2011 to correct
- bias is only in X-direction
- only affect 2-phase chips (50% of focal plane)
- bias is up to 150-200 mas
- this is *not*
 - CTE (wrong direction)
 - saturation (too early)



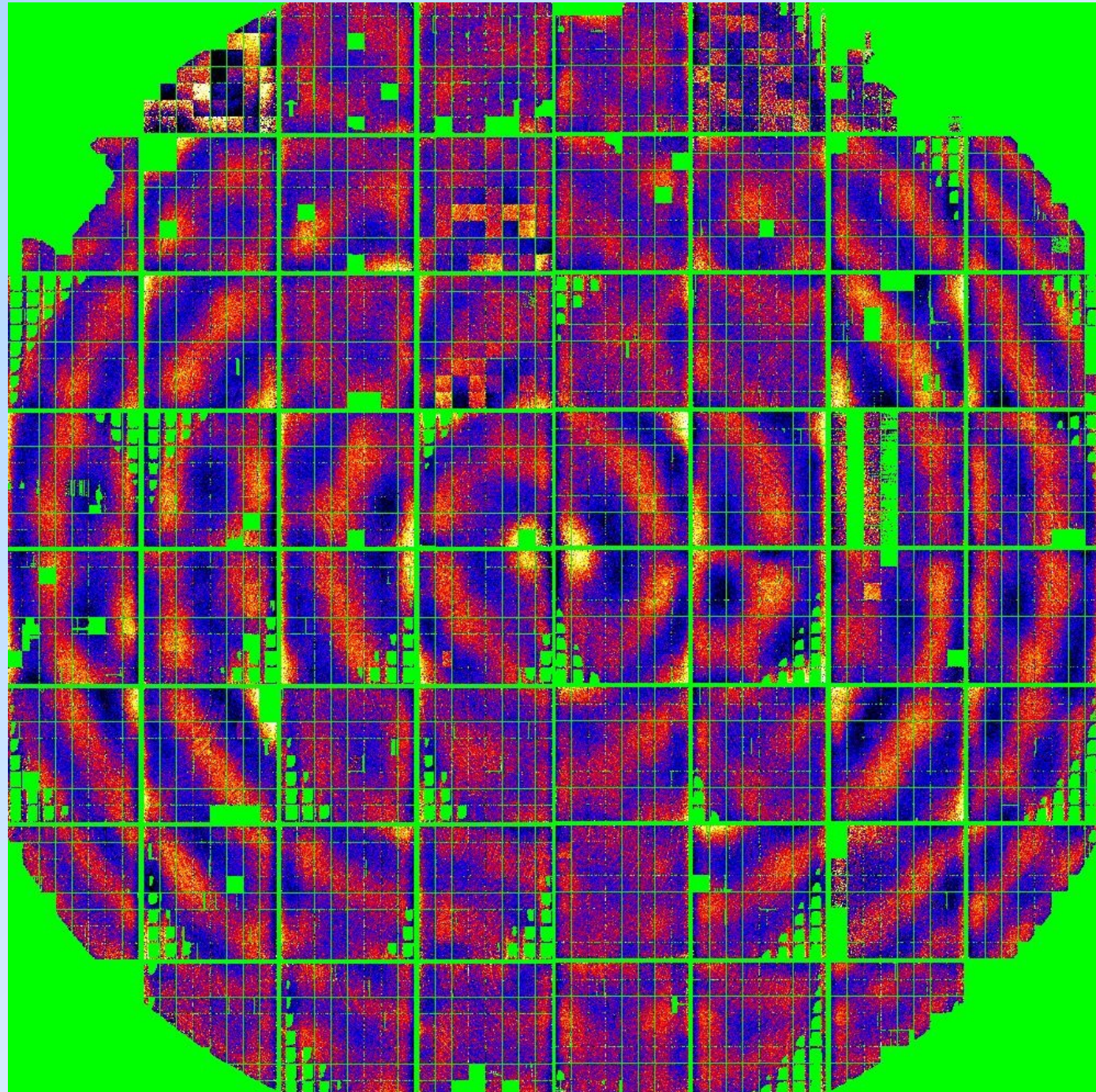
The Koppenhoefer Effect

- correction effect on typical errors (25mas \rightarrow 18mas)



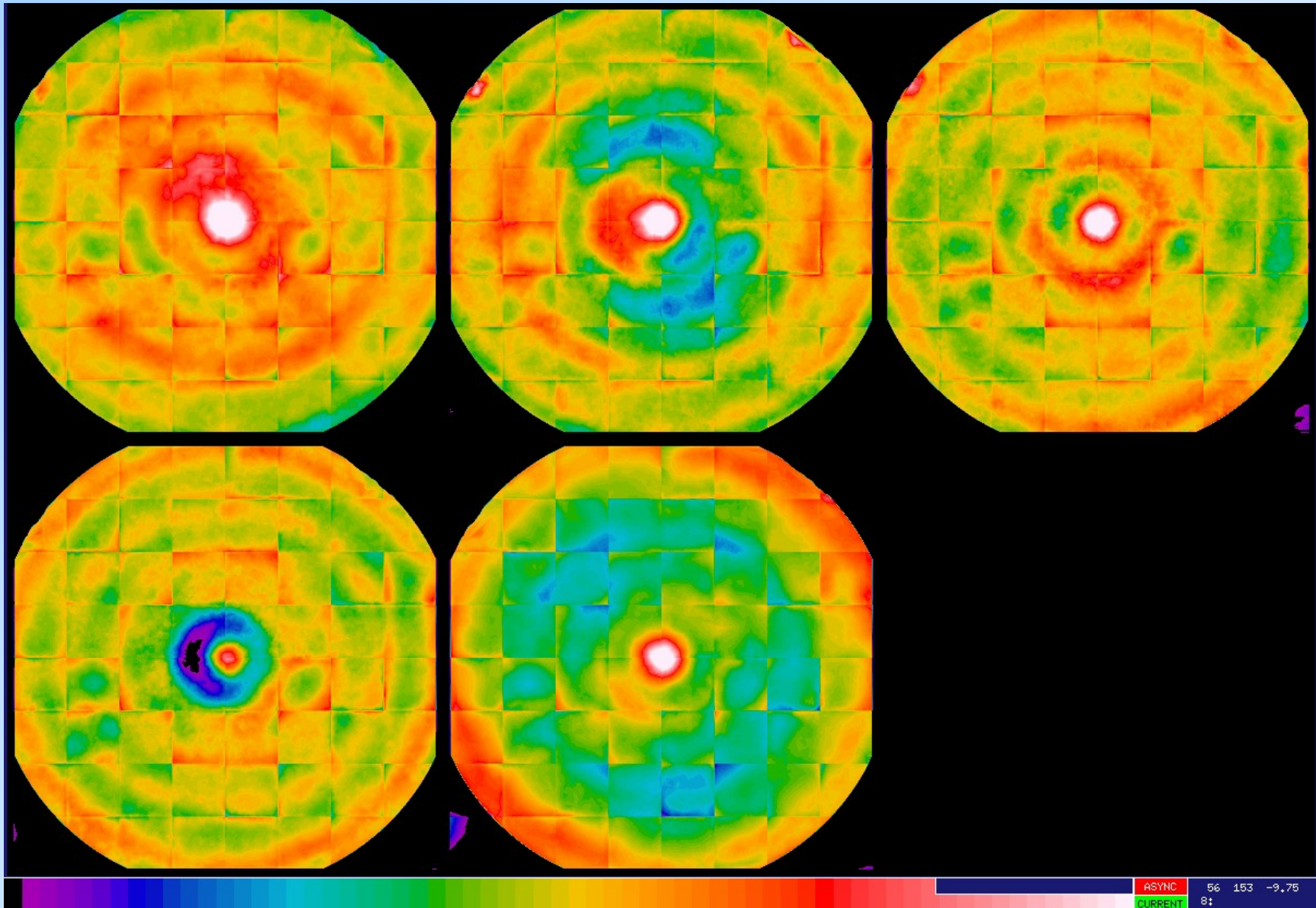
Astrometric Systematics

- mean residuals as a function of camera position
- 20 x 20 pixel bins
- i-band, dX shown
- large-scale structure similar to focal-plane deviations



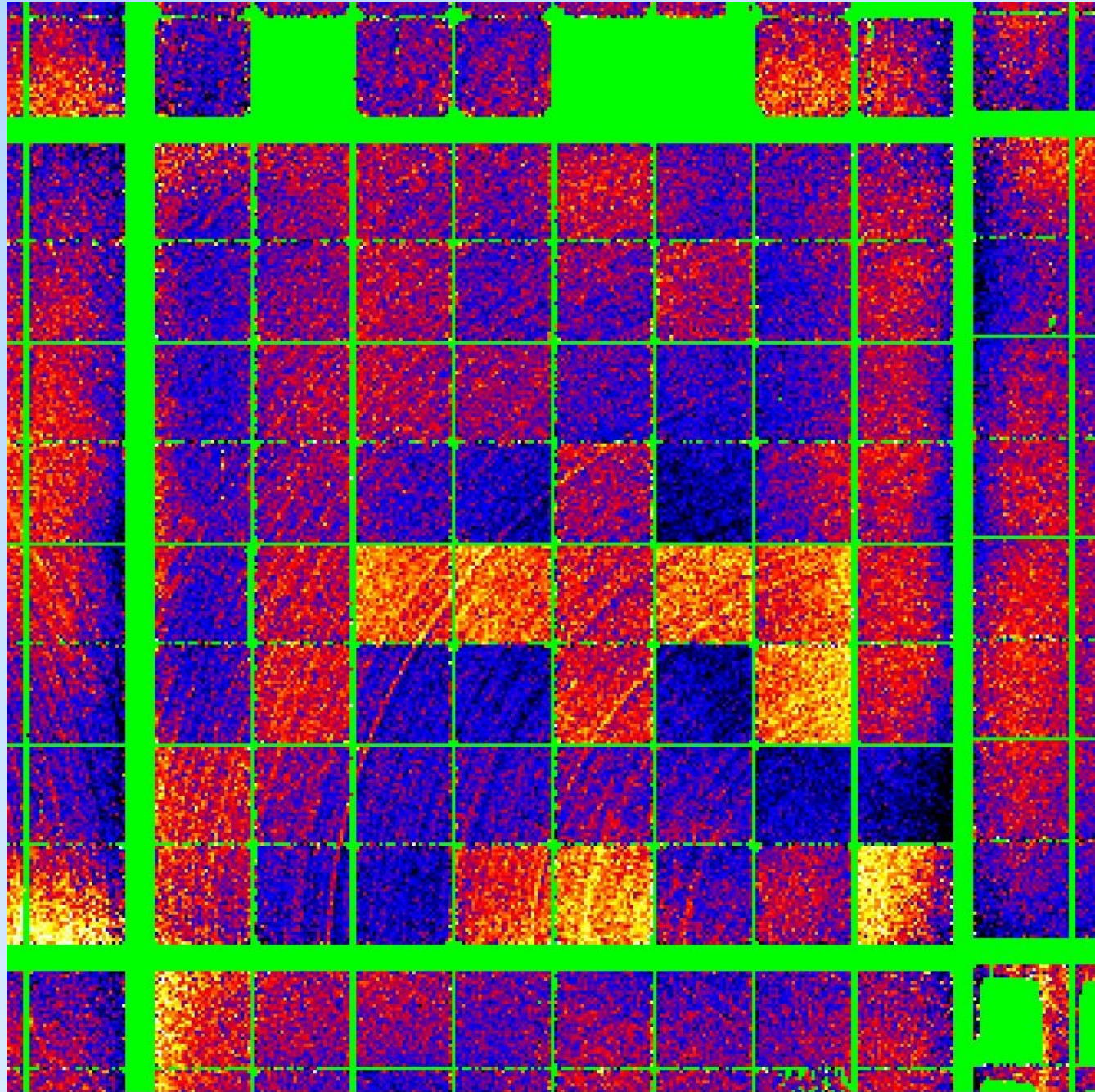
Astrometric Residuals : Large-Scale Circular Pattern

- Similar to focal-plane surface residual
- does the trend match in detail (e.g., per filter?)



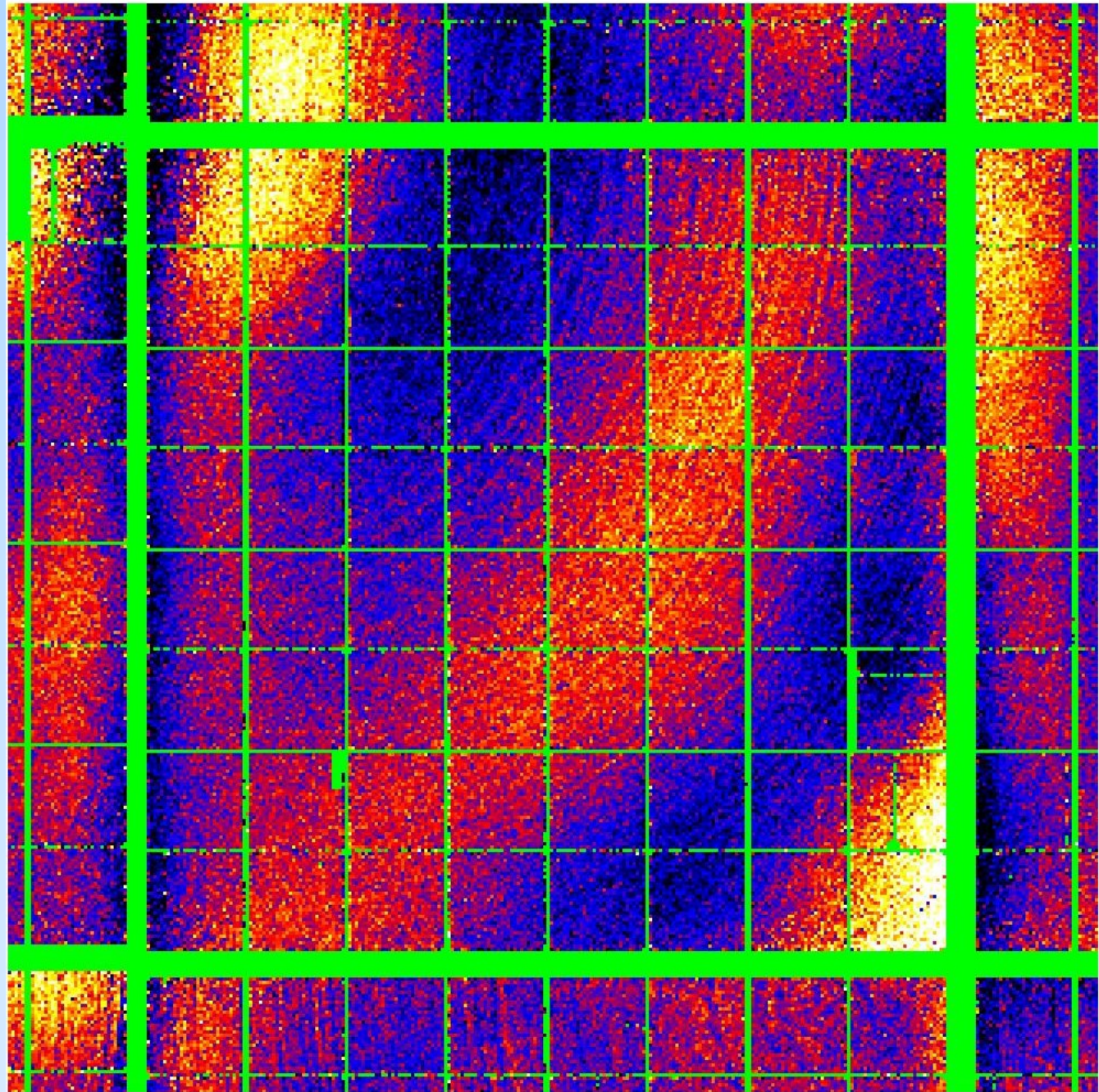
Astrometric Systematics : Cell-by-cell offsets

- mean residuals as a function of camera position
- other structure quantized per cell.



Astrometric Systematics : Tree Rings

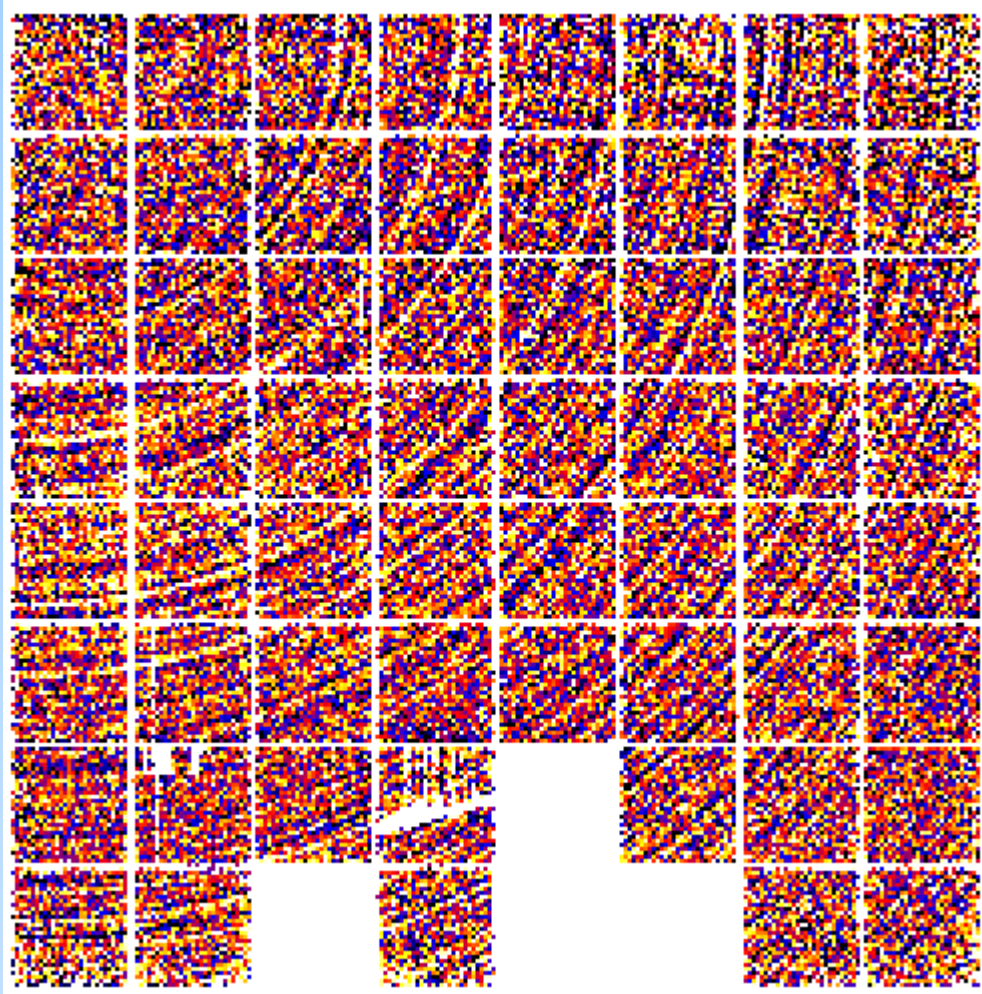
- mean residuals as a function of camera position
- 'tree-rings' (also seen by DES, others)



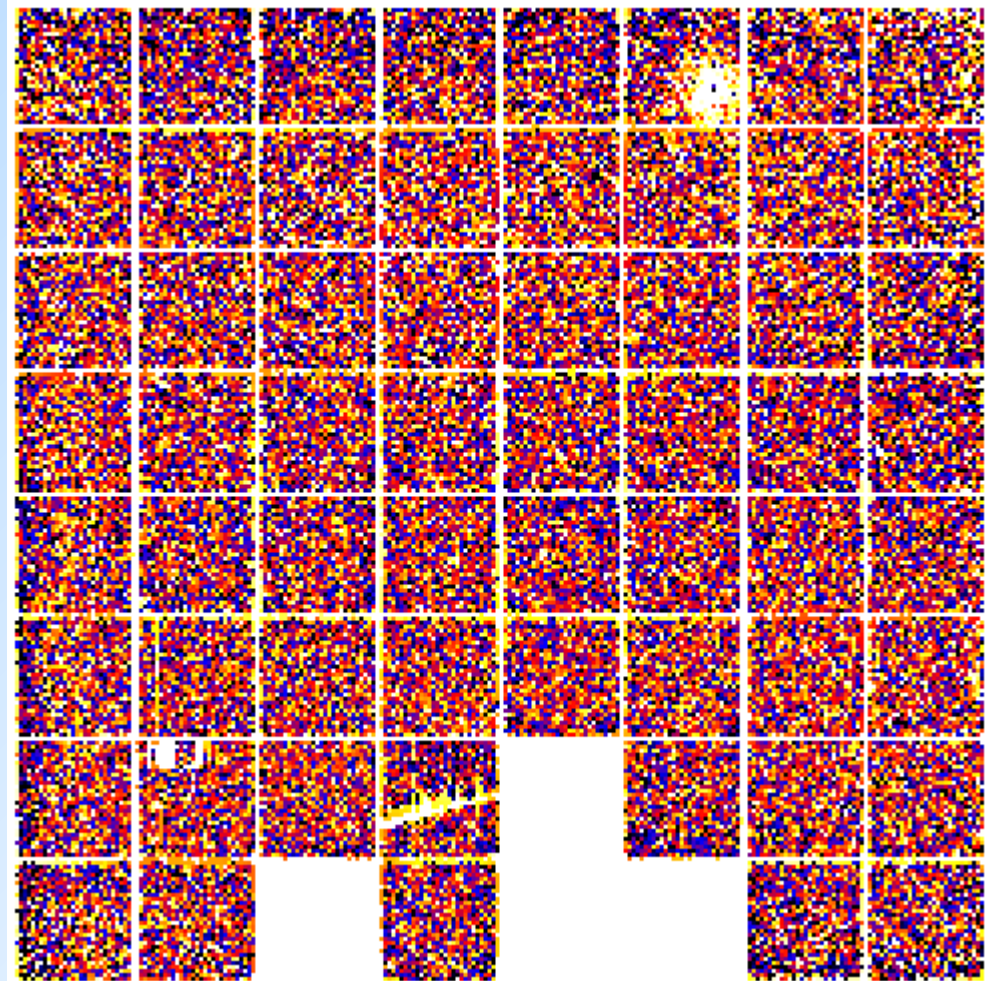
Astrometry Residual

- convert dX, dY to $dR, d\theta$
- subtract a smoothed version (high-pass filter)

radial component



tangential component



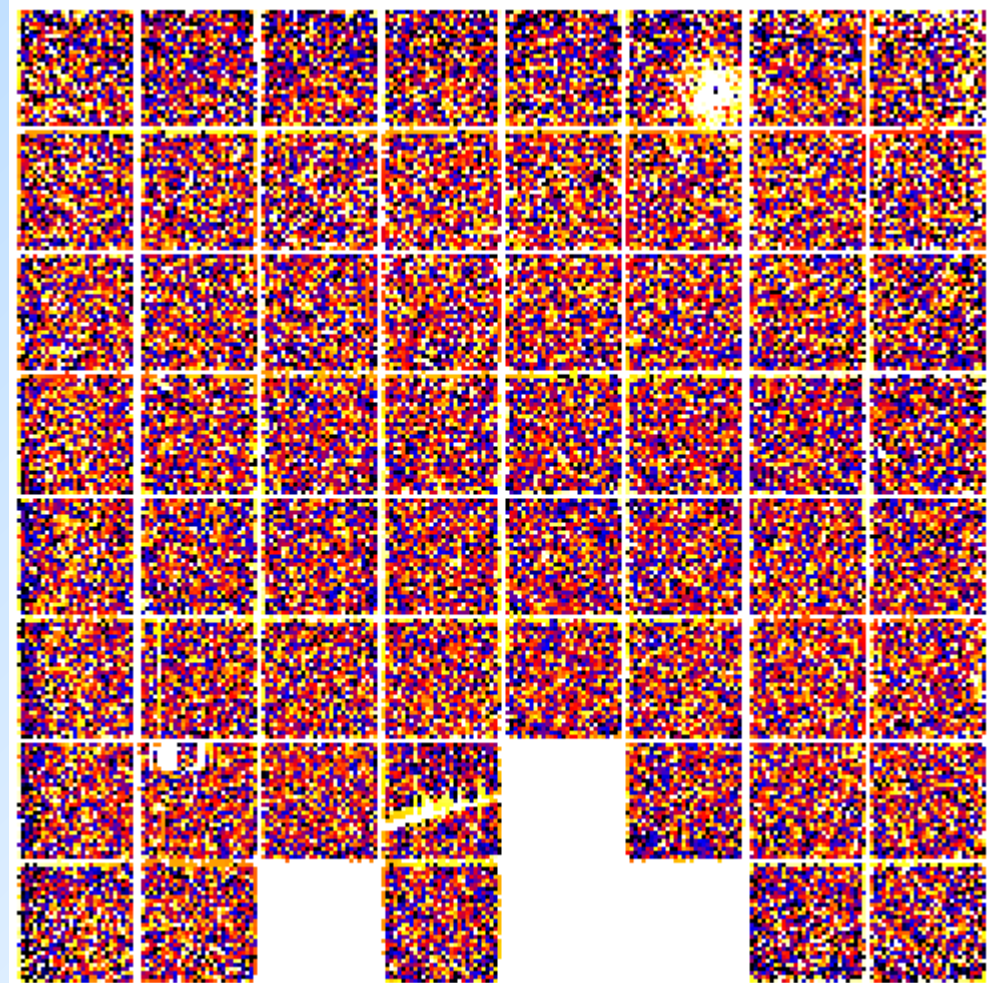
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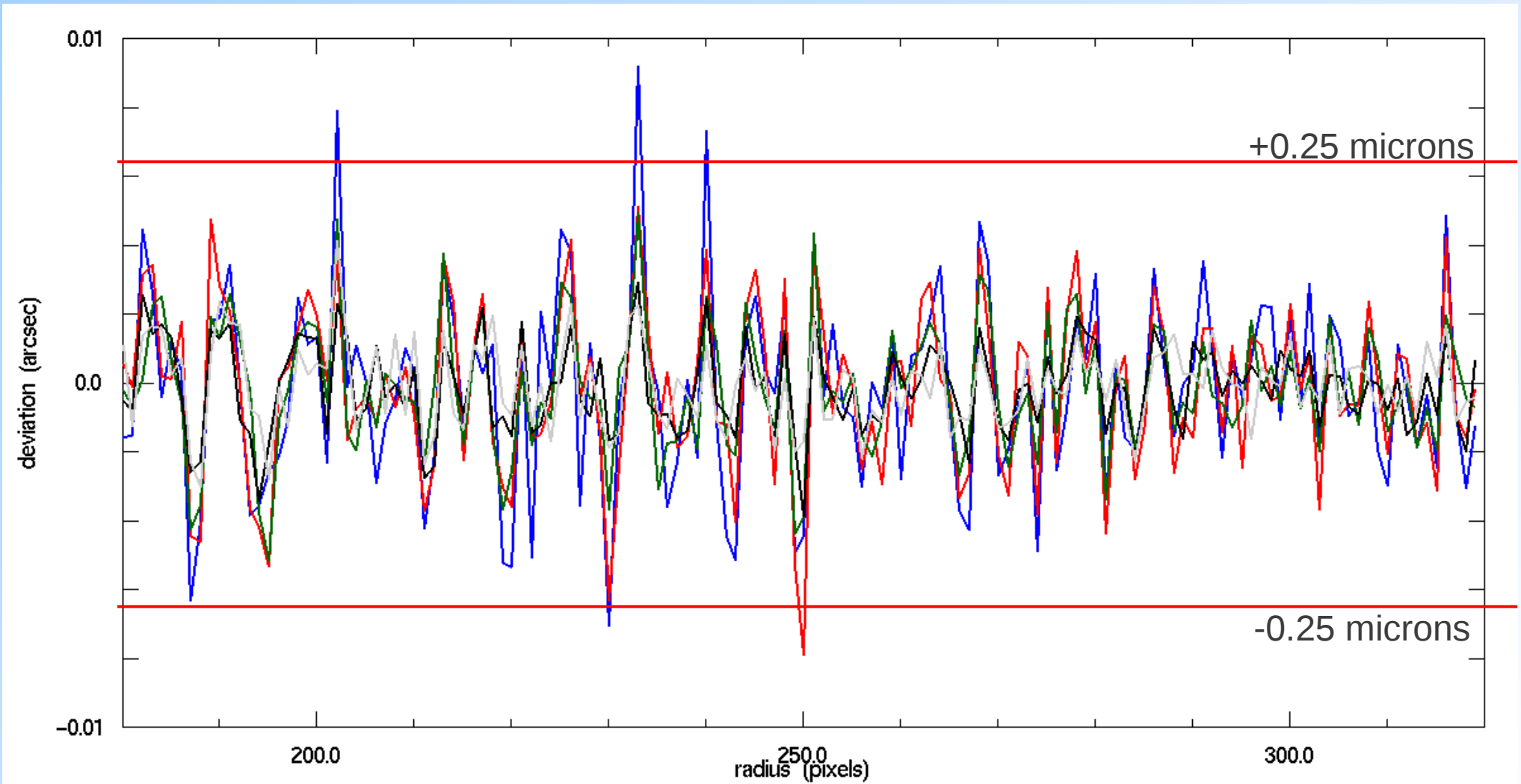


tangential component



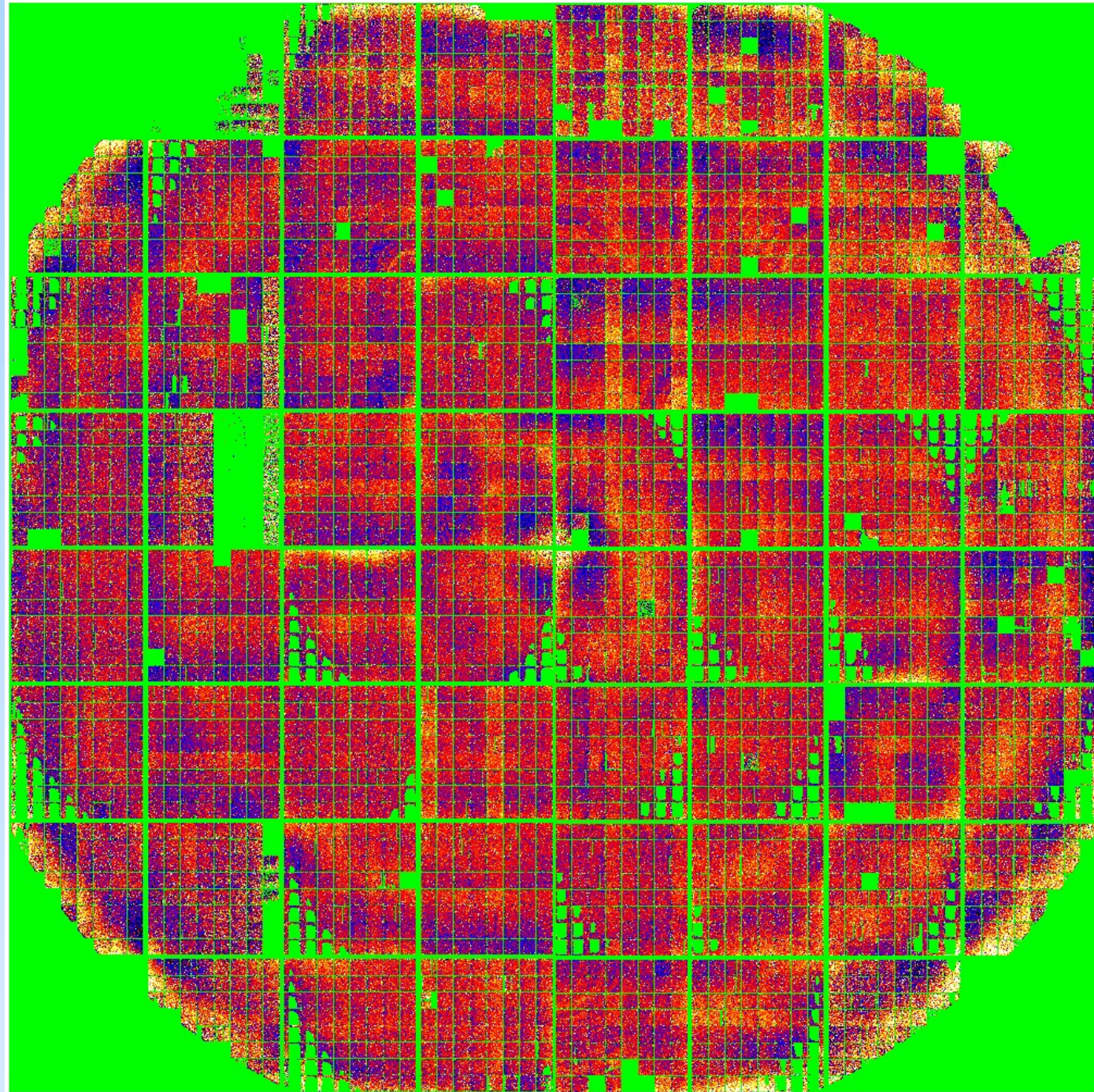
Astrometric Deviations

- *grizy* deviations are correlated, with scaling:
- $dR_f / dR_g = (0.50, 0.36, 0.23, 0.21)$ for $f = (r, i, z, y)$
- *(grizy)* = (blue, red, green, black, grey)



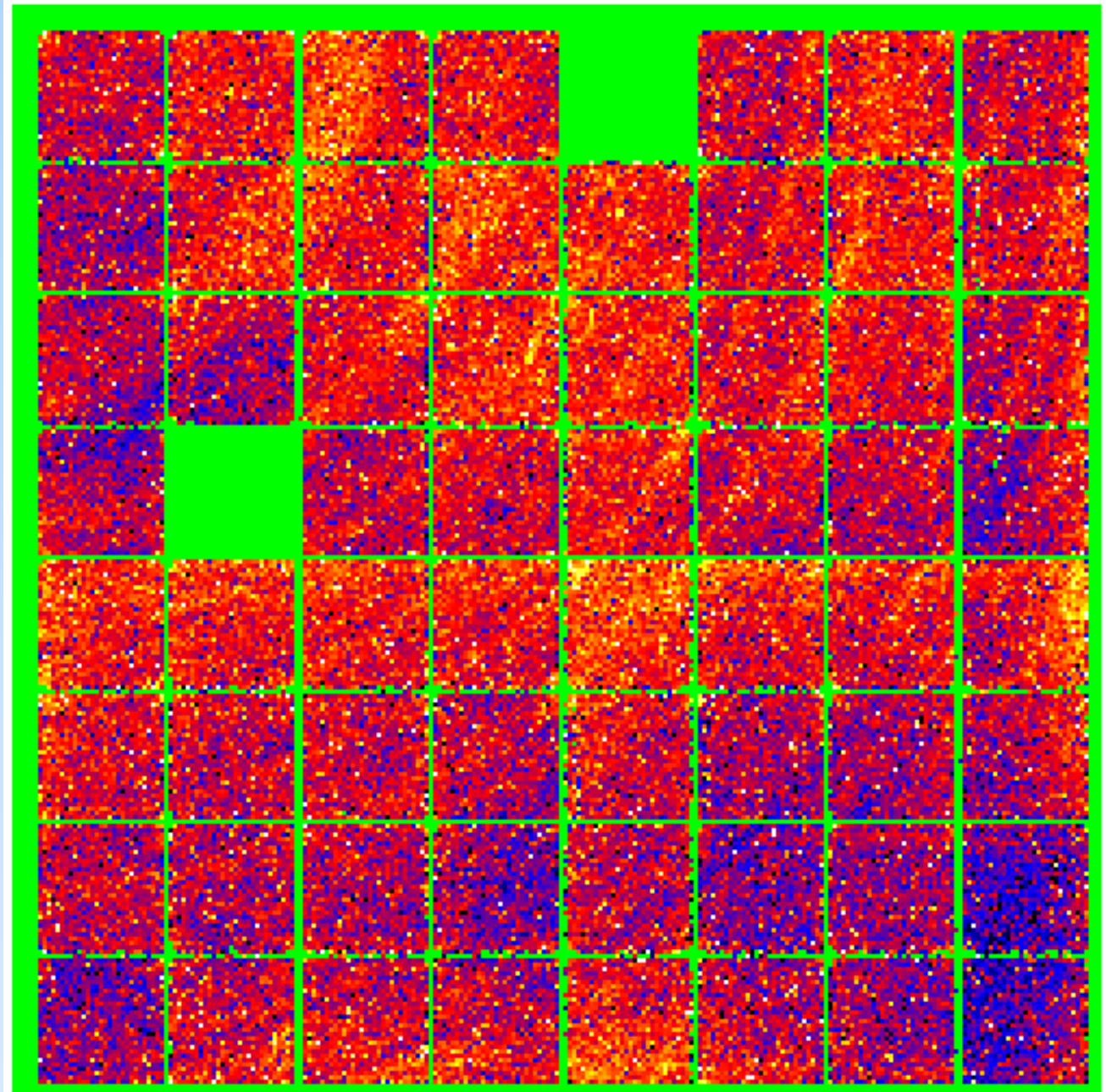
Photometric Systematics : per-exposure residuals

- mean residuals as a function of camera position
- 20 x 20 pixel bins
- i-band shown
- central tent
- residual of 2x2 flats



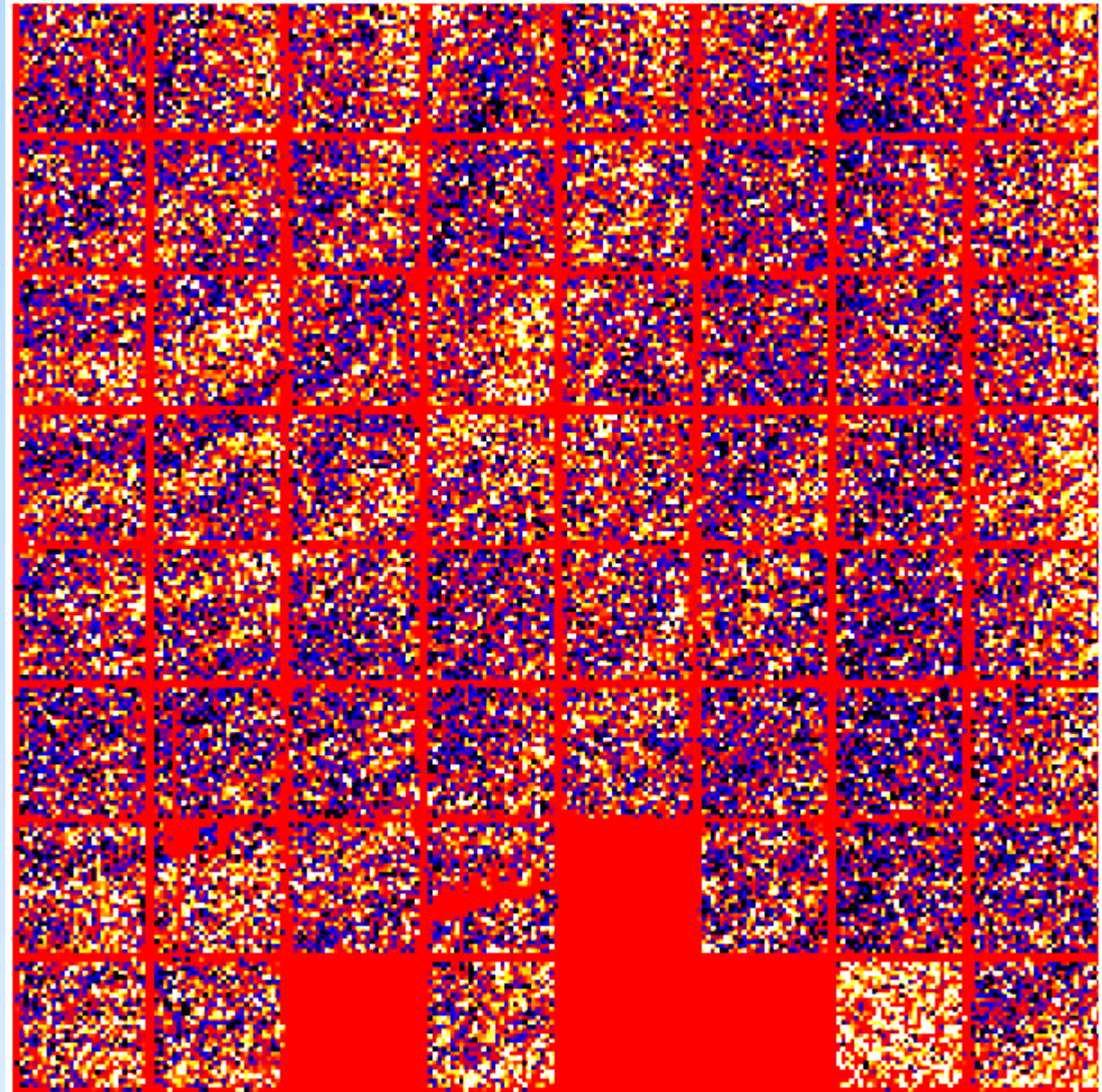
Photometric Systematics : per-exposure residuals

- mean residuals as a function of camera position
- 20 x 20 pixel bins
- i-band shown
- central tent
- residual of 2x2 flats
- 'tree-rings' also seen



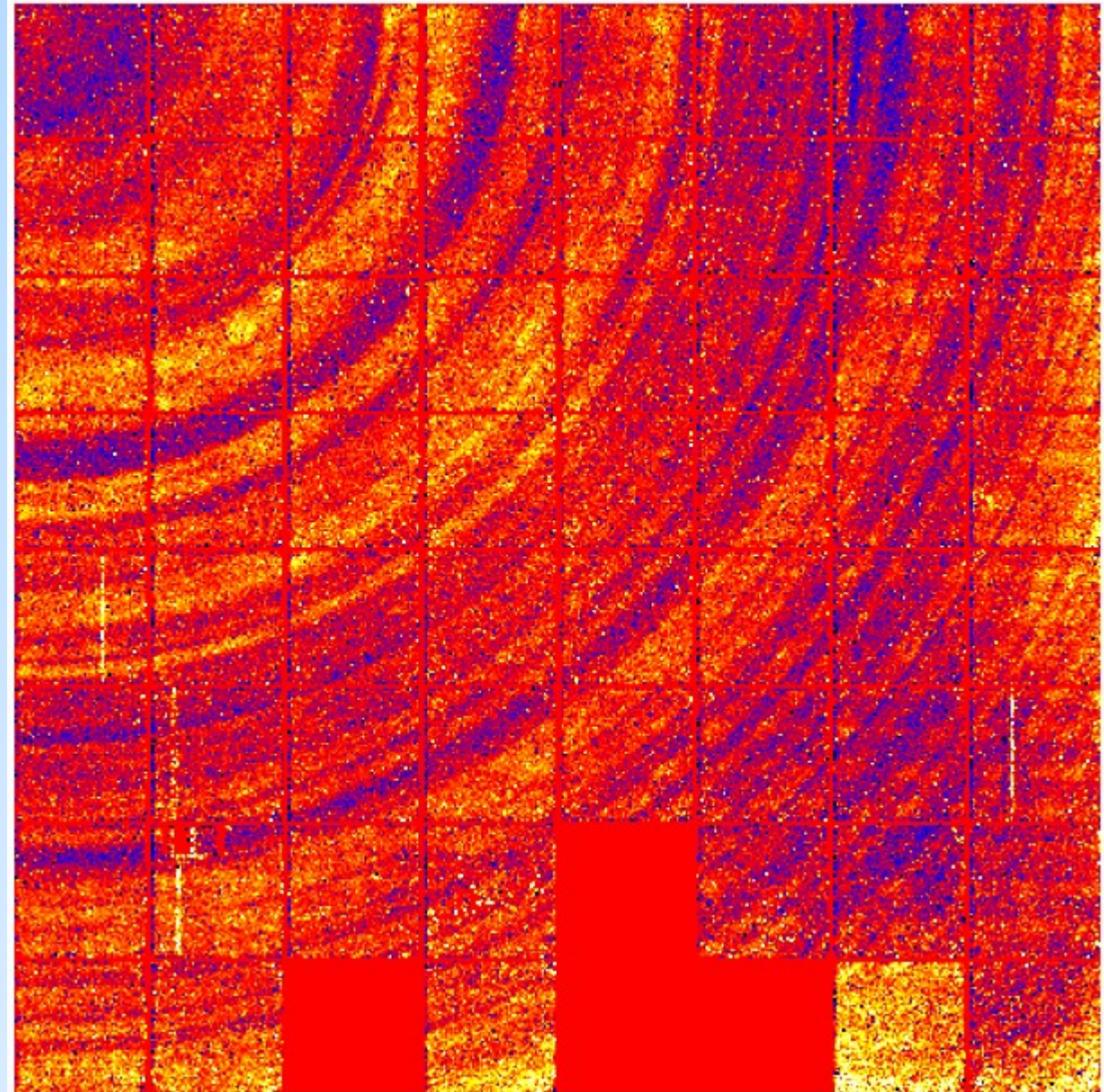
Photometric Residuals : Sample Chip

- Photometric Residuals (3pi)
- Similar to tree rings
- S/N is not great



Photometric Residuals : Sample Chip

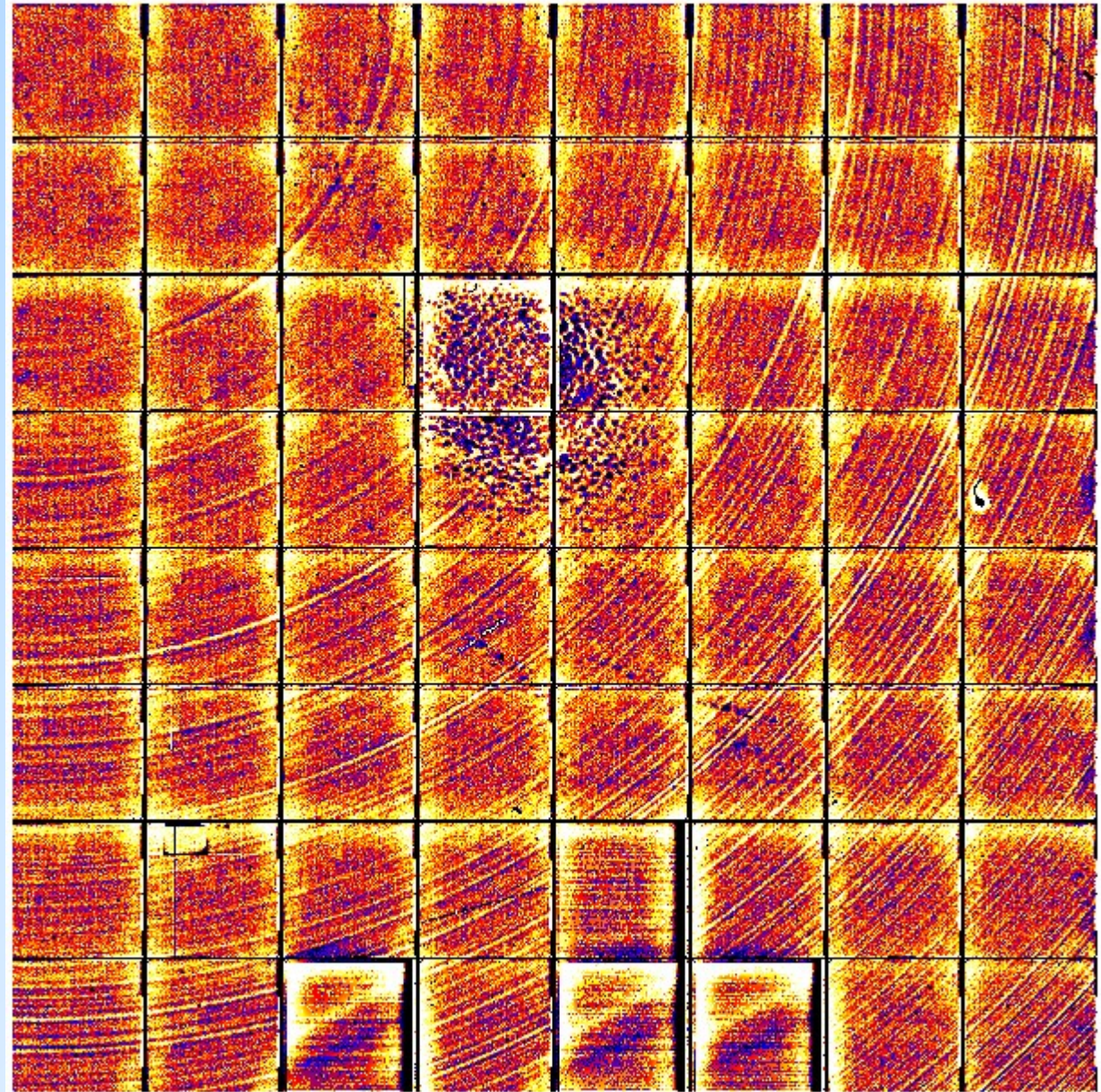
- Photometric Residuals (MD)
- From D. Finkbeiner
- Pattern matches 3pi
- S/N is much higher
- Is this a Jacobian Effect?
 - astrometric effect moves stars but squeezes flat-field light
 - residuals are imprint of over/under correction in the flat field?



Photometric Residuals : Sample Chip

- Flat-field shows the rings

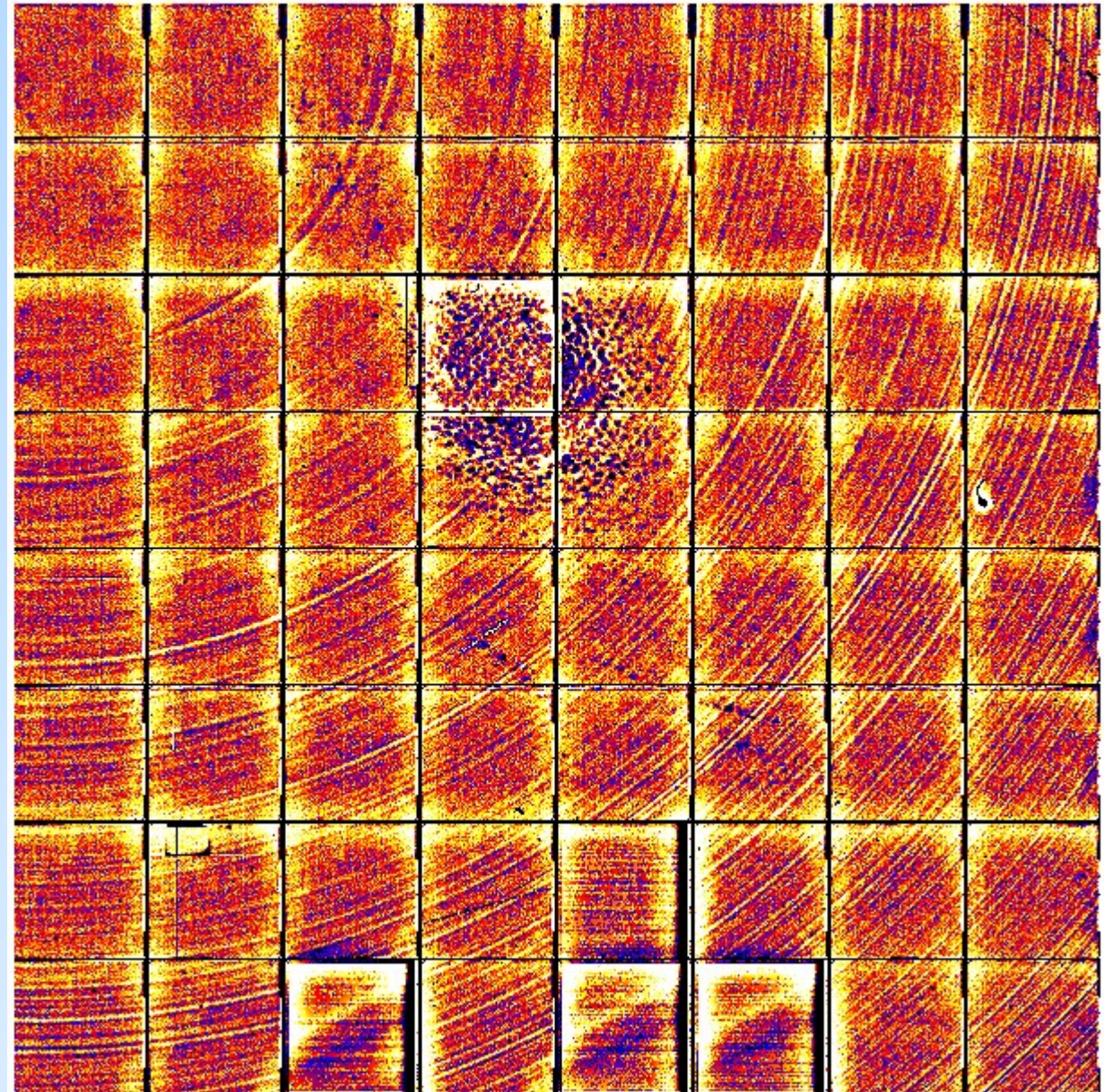
monochromatic flat @ 630nm



Photometric Residuals : Sample Chip

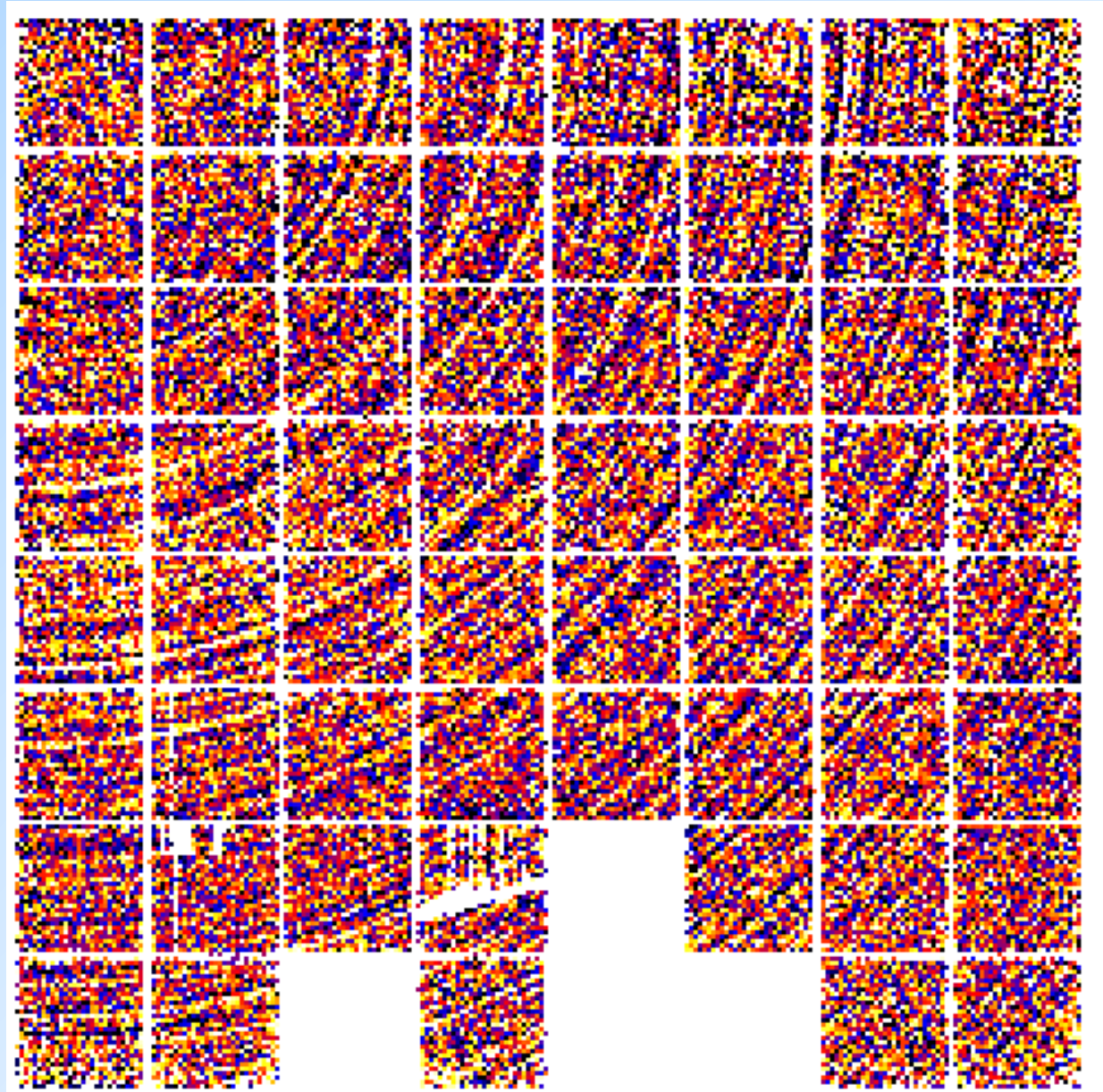
- Flat-field shows the rings
- **But wait!**
 - **flat field looks like astrometric residuals**

monochromatic flat @ 630nm



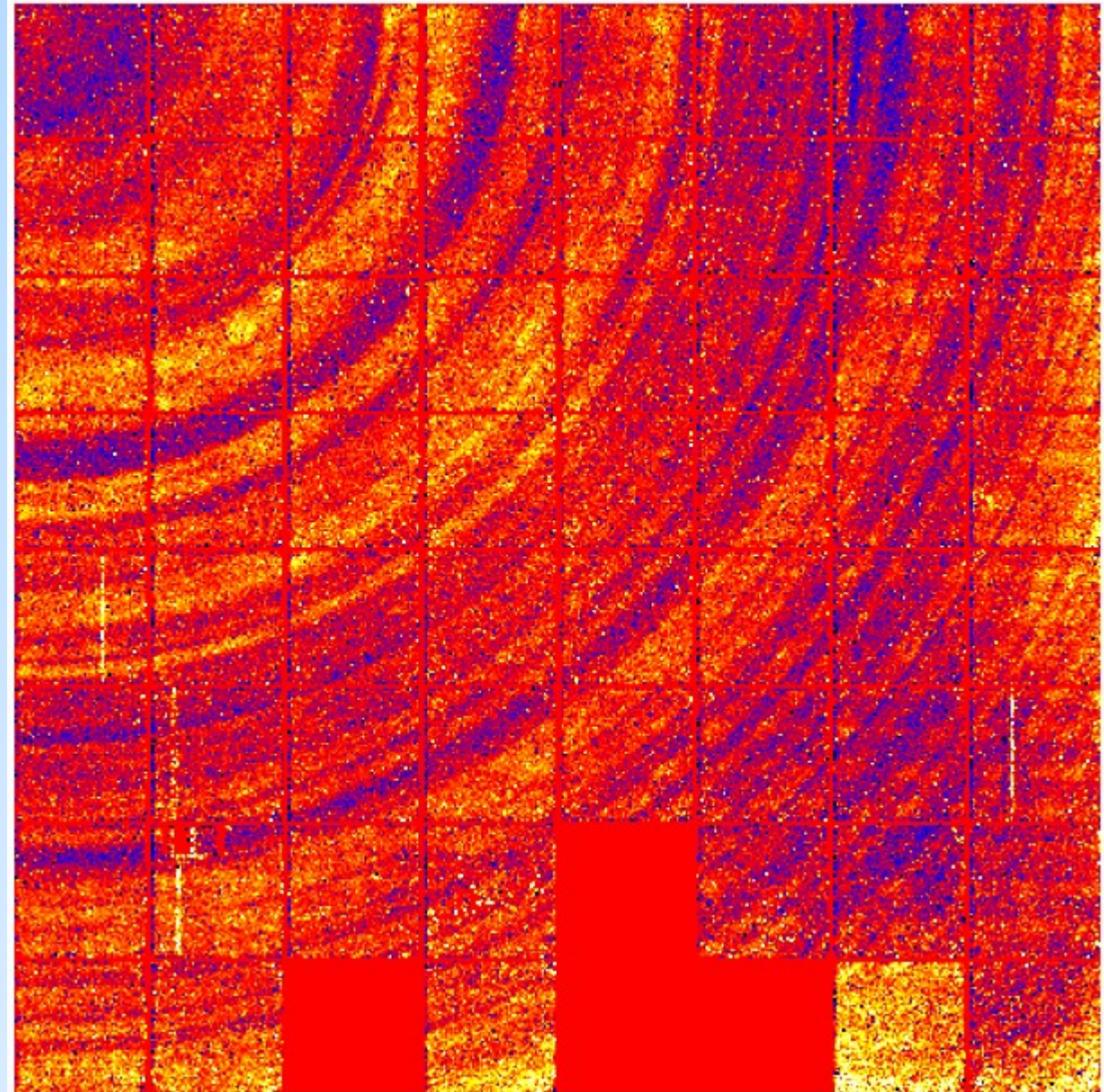
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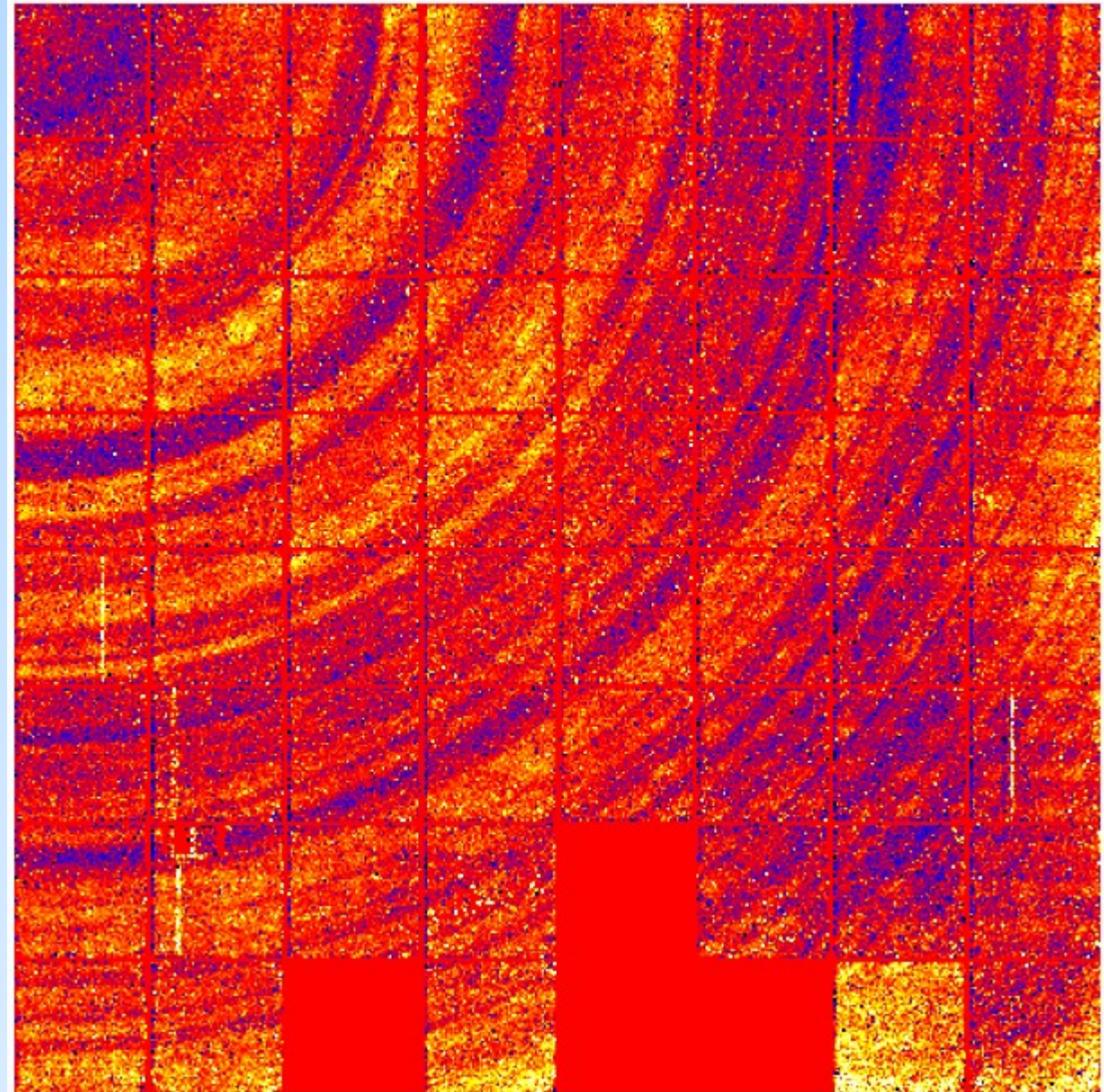
Photometric Residuals : Sample Chip

- Flat-field shows the rings
- **But wait!**
 - **flat field looks like astrometric residuals**
 - **not like photometric residuals**



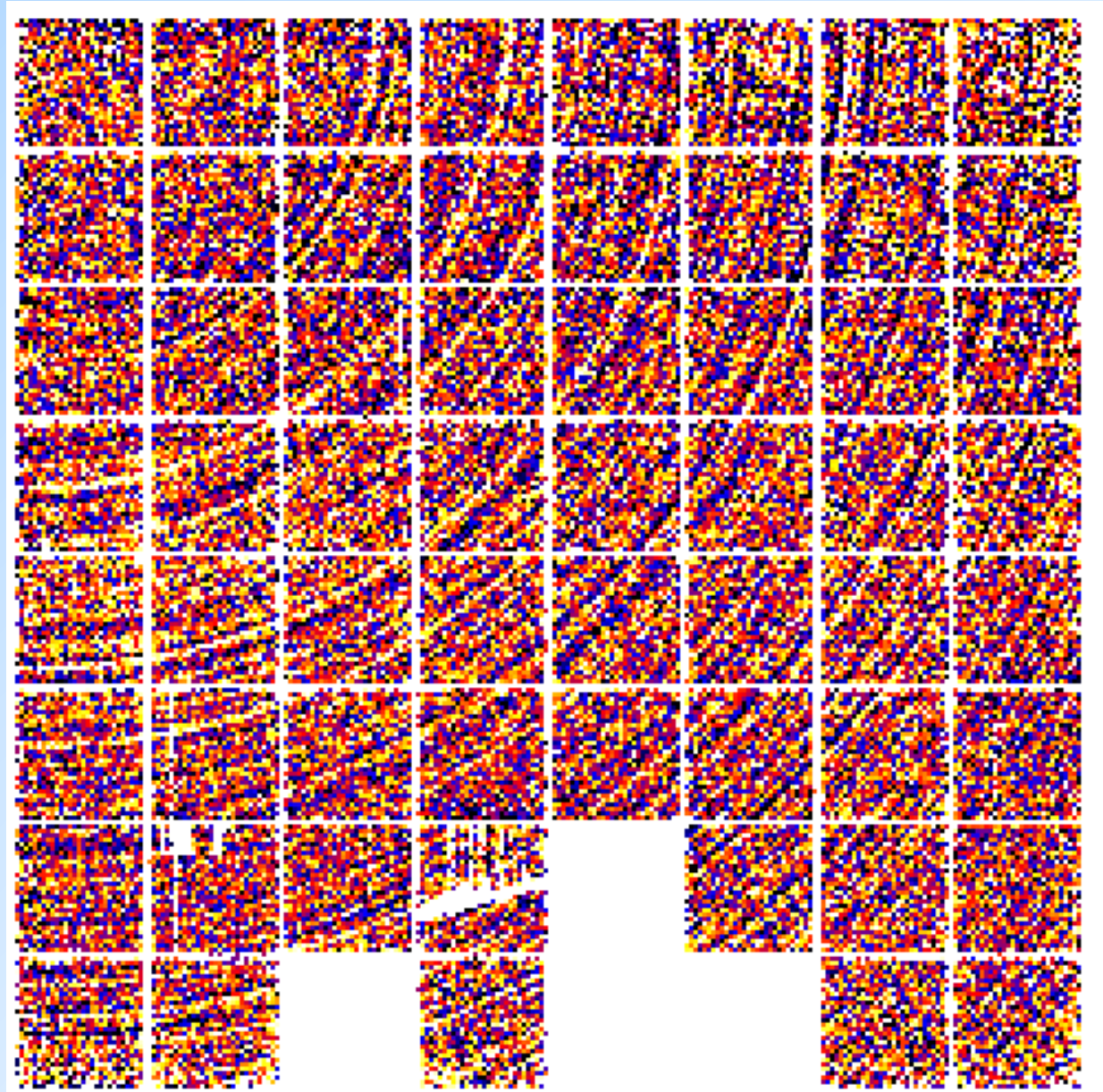
Photometric Residuals : Sample Chip

- Flat-field shows the rings
- **But wait!**
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 - not like photometric residuals
- Also we see:
 - $dR \sim \text{grad } dM$
- But Jacobian effect wants
 - $dM \sim dR$



Photometric Residuals : Sample Chip

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- **But wait!**
 - flat field looks like astrometric residuals
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- Also we see:
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Summary

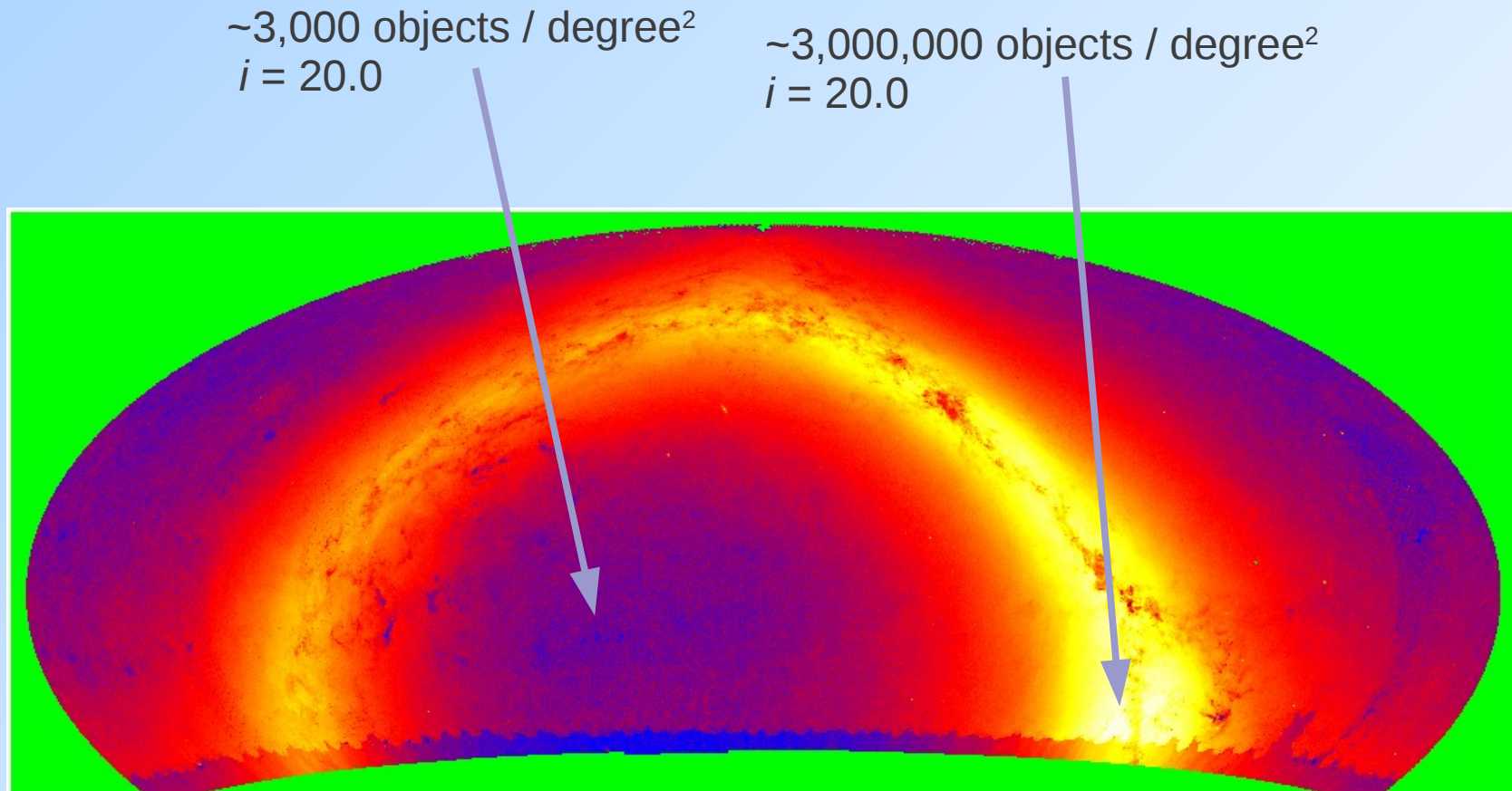
- PS1 is achieving good astrometry & photometry
- There is room for improvement
 - static systematics
 - finer spatial modeling of PSF variations
 - finer spatial modeling of astrometric corrections
 - stellar density is the ultimate limiting factor
- Tree-rings show up in astrometry and photometry
- We do not really understand the tree rings...

Backup Slides Follow:

- Astrometry and the atmosphere
- Spatial Sampling

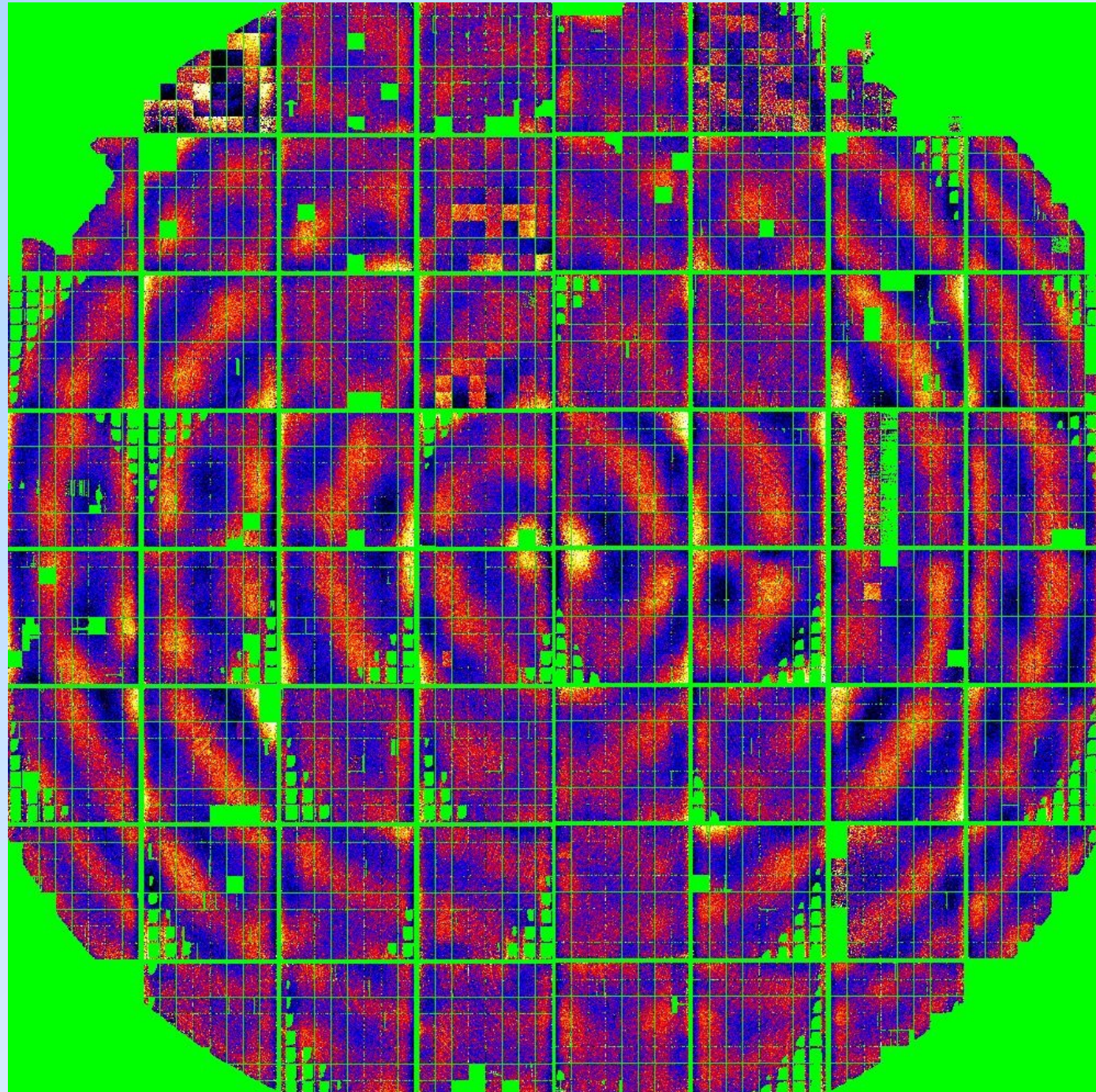
Finer Sampling

- PSF modeling and astrometric correction need references
- Limit of spatial sampling is stellar density
 - $1000 \text{ deg}^{-2} \rightarrow \sim 6 \text{ arcmin}$
 - $10,000 \text{ deg}^{-2} \rightarrow \sim 2 \text{ arcmin}$



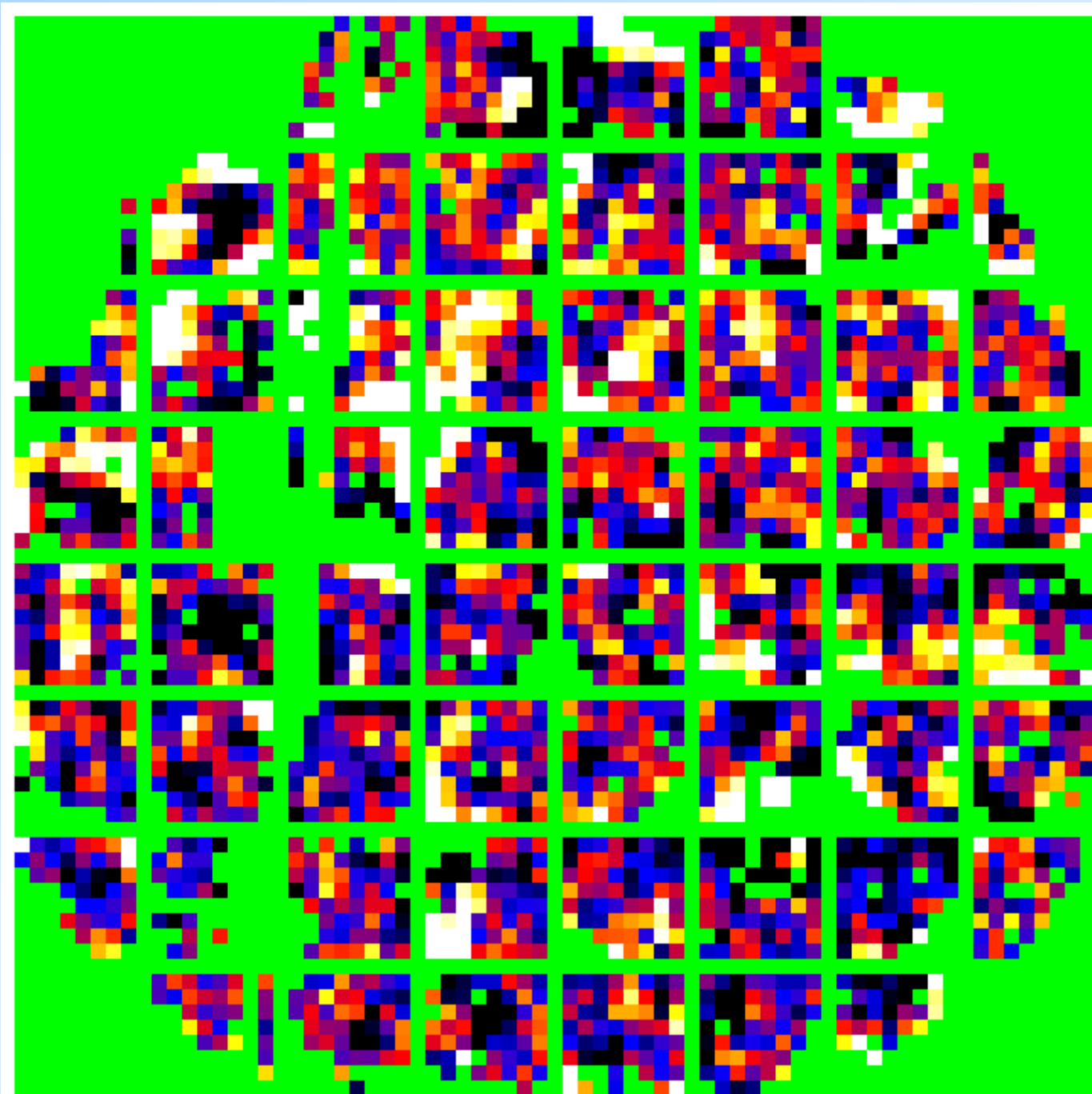
Astrometric Systematics

- mean residuals as a function of camera position
- constant systematics contribute ~ 5 mas
- **overall systematics are ~ 20 mas**
- **what is the source?**



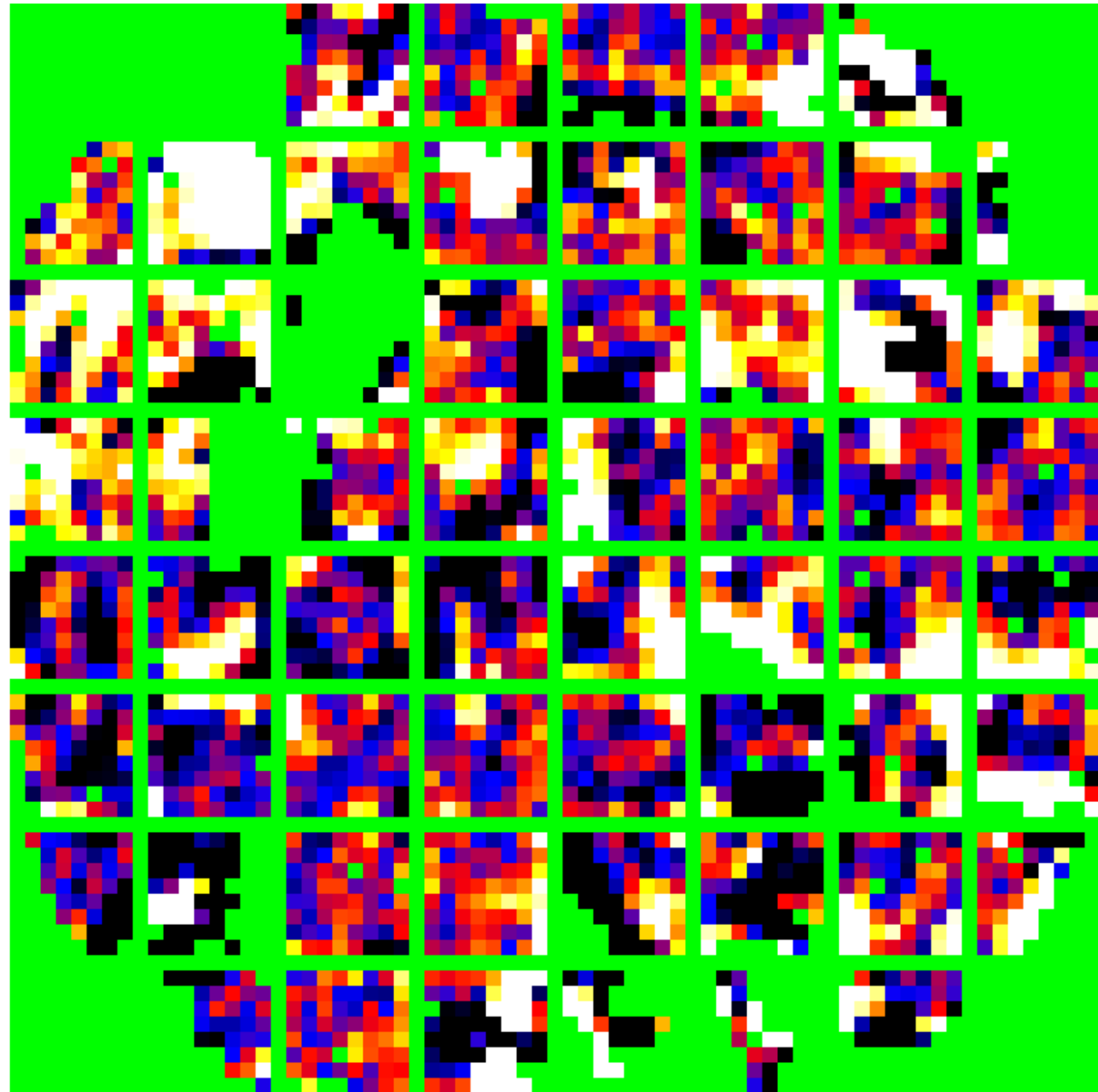
Astrometric Systematics : per-exposure analysis

- choose dense fields
- find mean astrometric offsets in cells for each exposure
- we have dX and dY astrometric offsets



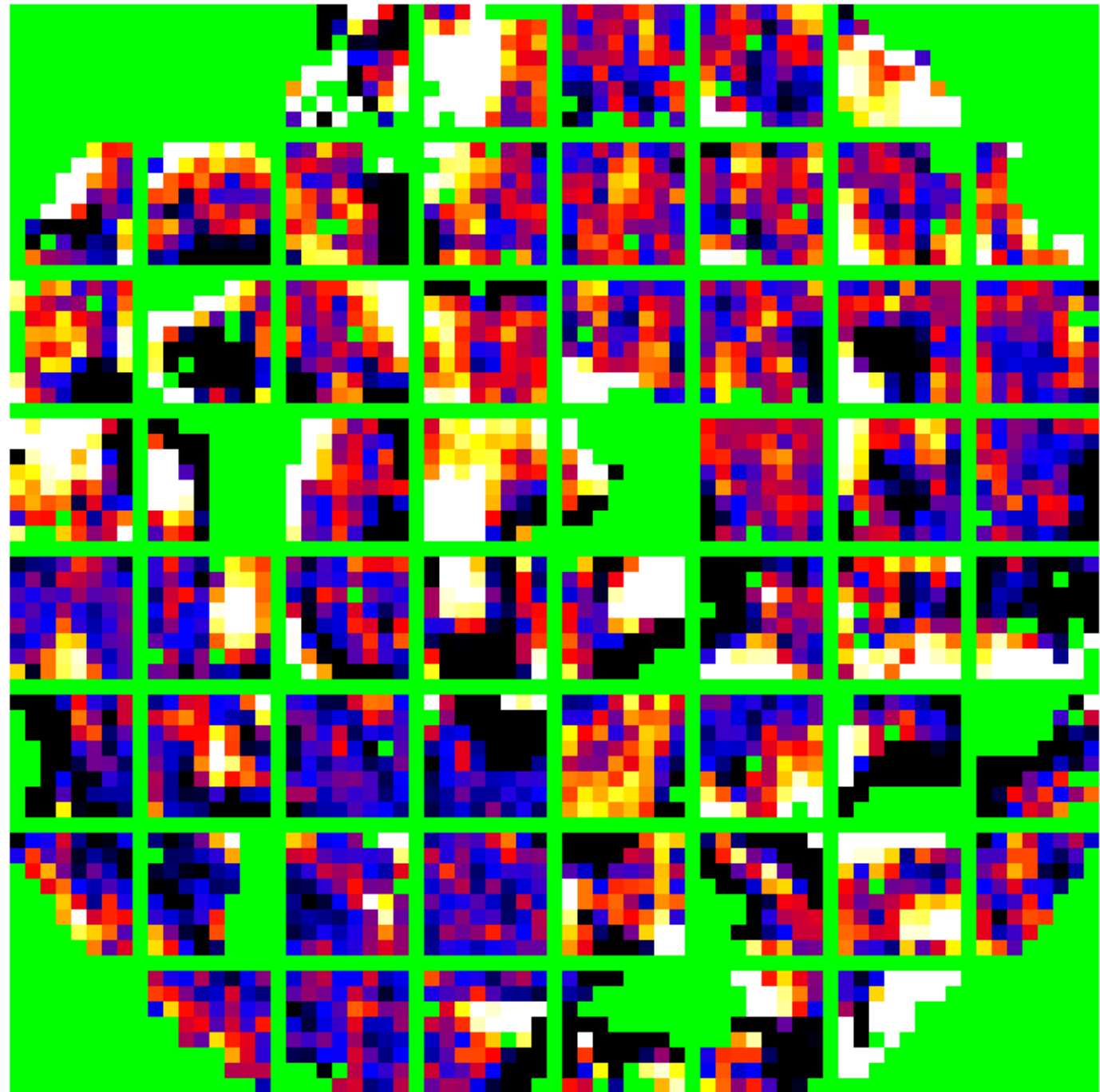
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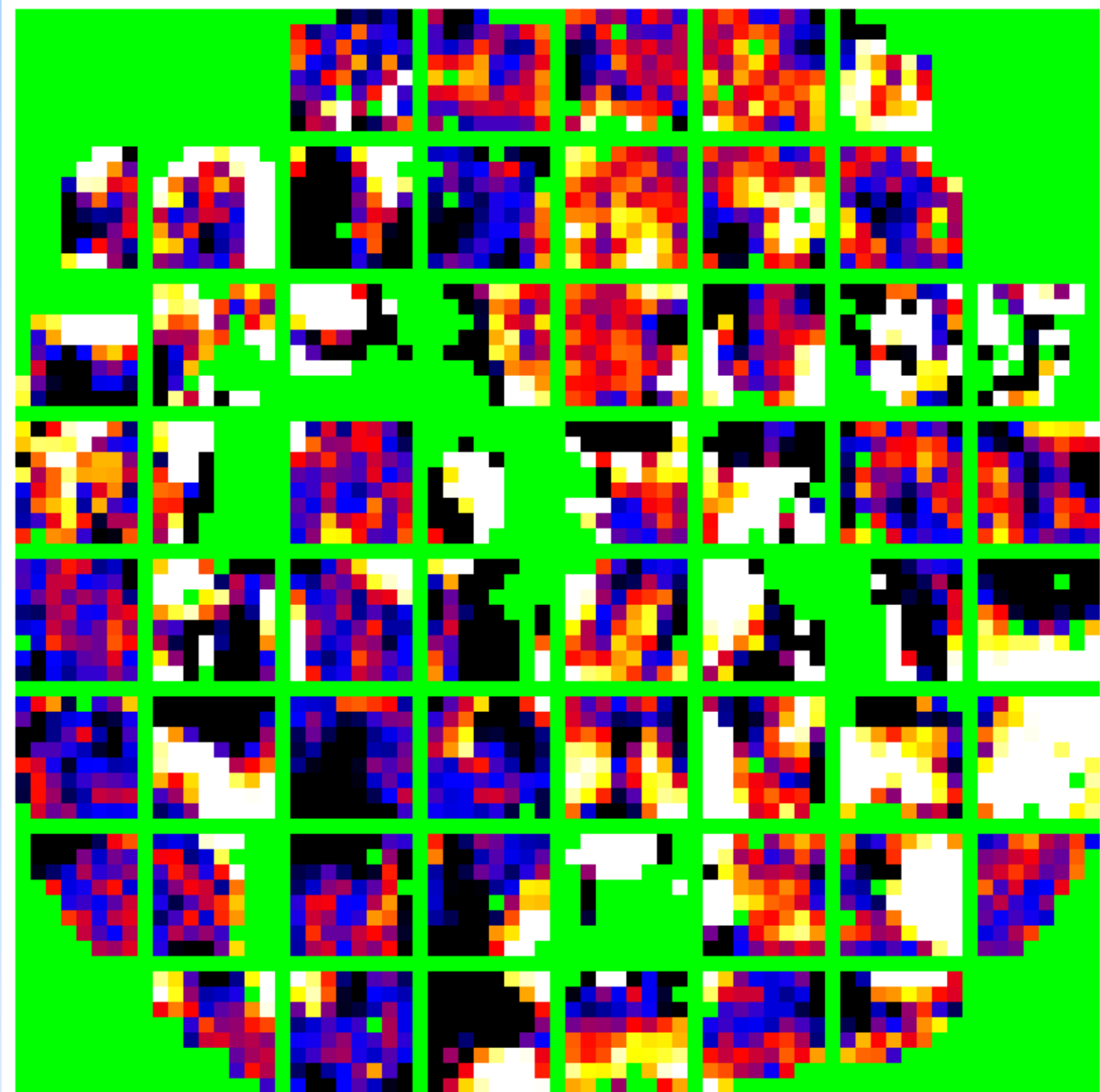
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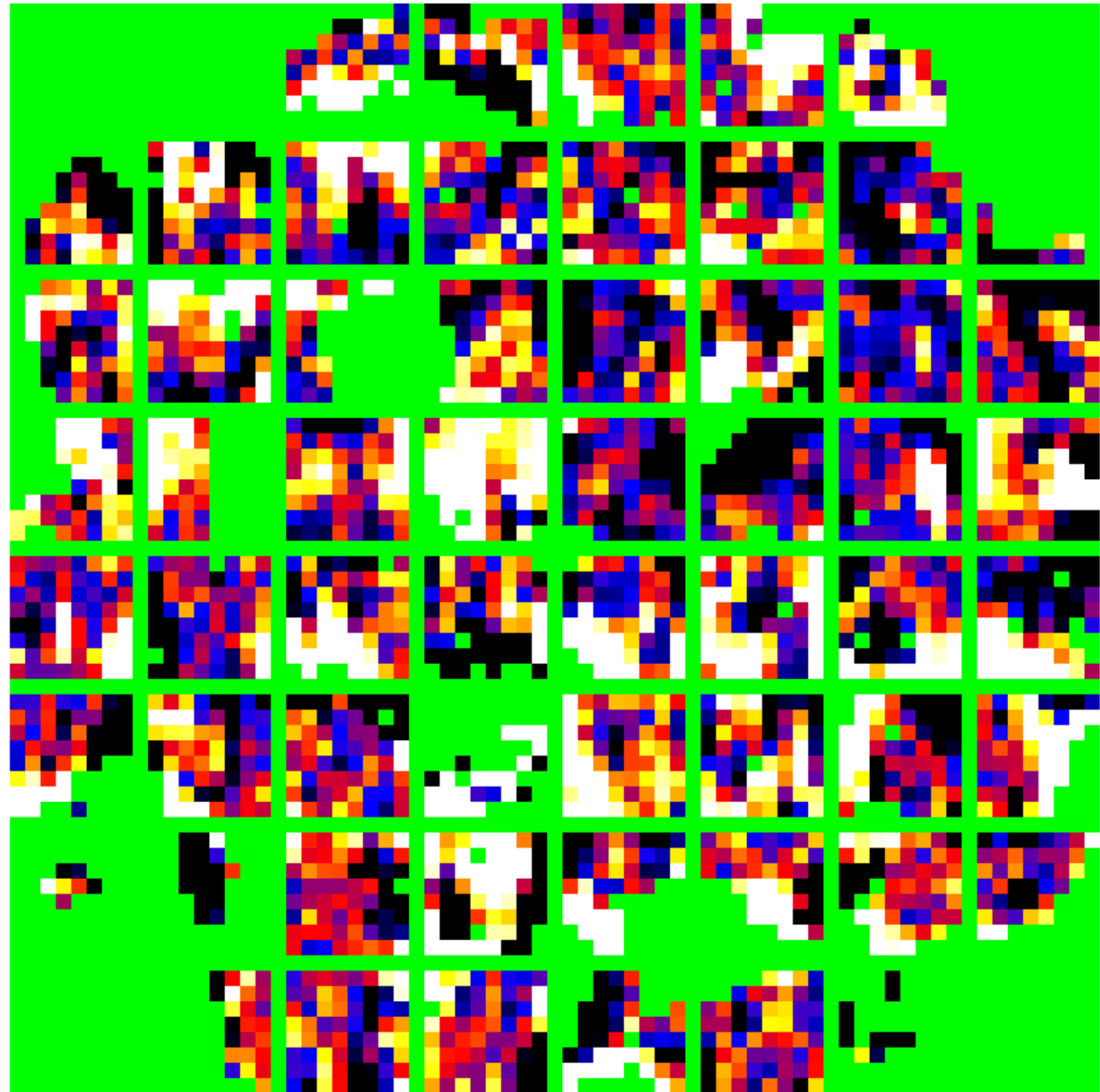
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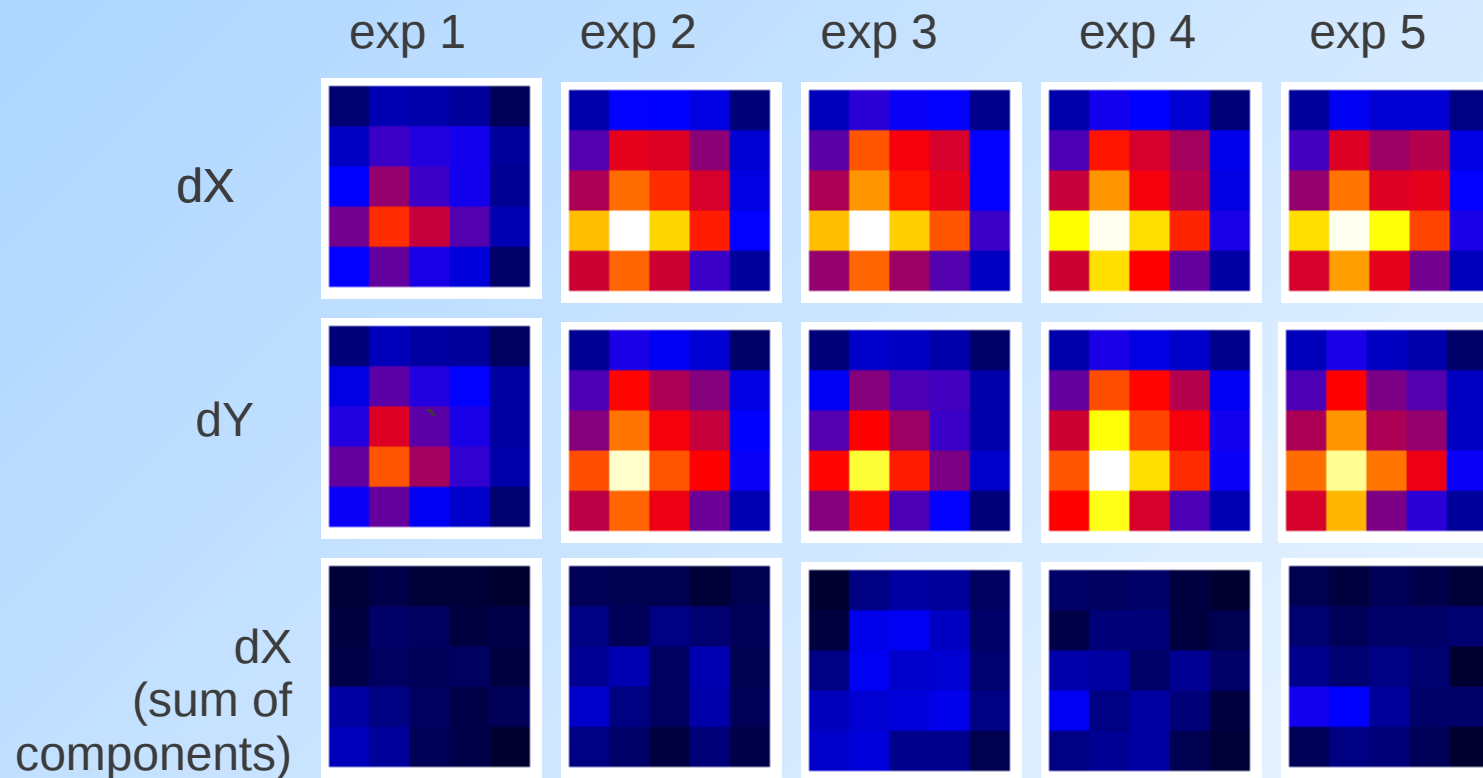
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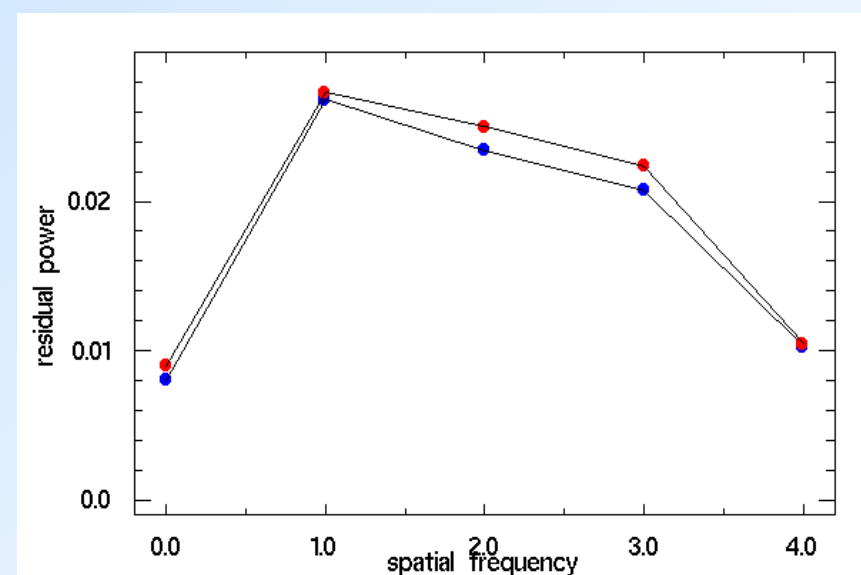
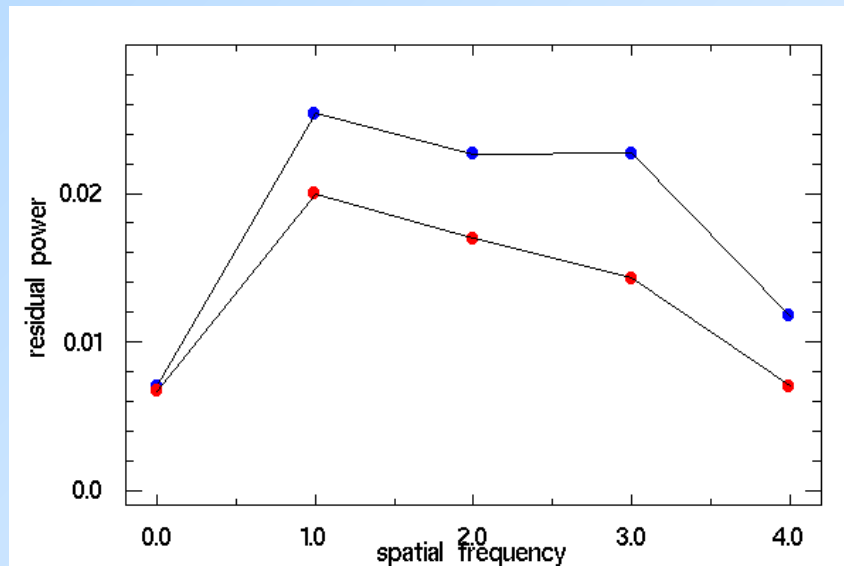
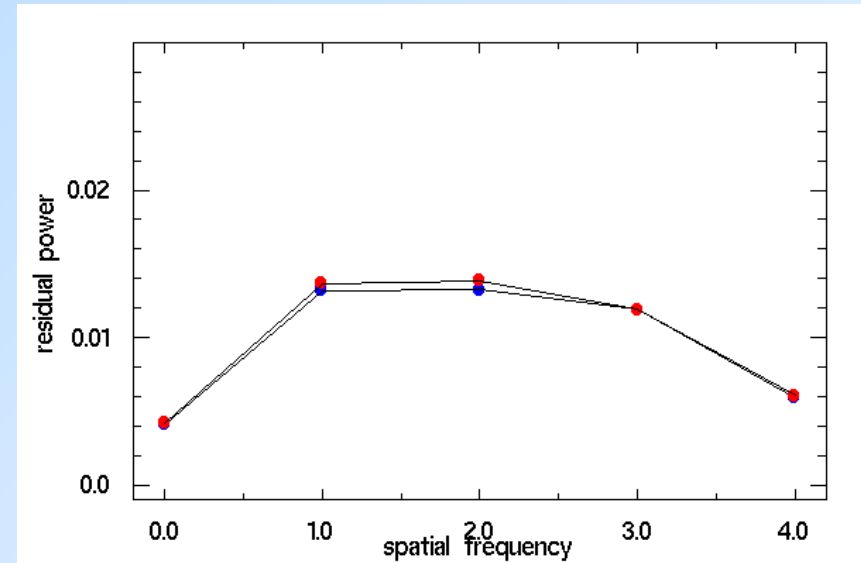
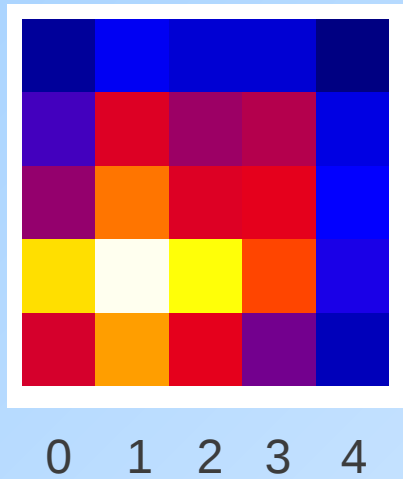
Astrometric Systematics : FFT of per-exposure data

- FFT of data per chip
- sum of power
- not coherent over camera



Astrometric Systematics : FFT of per-exposure data

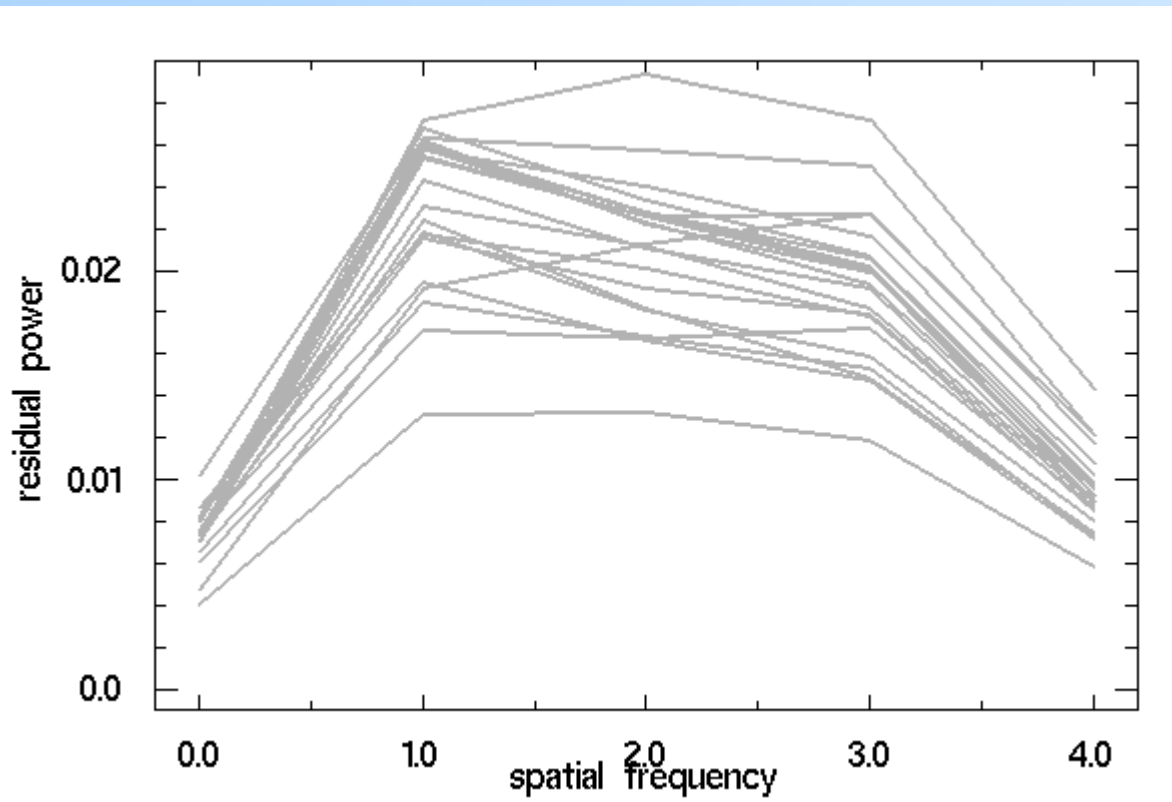
- sum the power (in quadrature) in annuli
 - residual power (N) = $\text{sum}(F \leq N) - \text{sum}(F < N)$



Astrometric Systematics : Significant Spatial Frequencies

- power distribution varies from exposure to exposure
- **FFT interpretation of sigma seems OK**
- **correction at 4x4 or 6x6 cells per chip would improve model**
- **correction at 8x8 cells yields diminishing returns**
- **correction @ 4x4 requires 160 stars / chip (~ 1000 / deg^2)**
- **correction @ 6x6 requires 360 stars / chip (~ 3000 / deg^2)**

residual power for 21 exposures



camera vs implied sigma

